



# The MicroBooNE Experiment

On behalf the MicroBooNE  
collaboration

Raquel Castillo Fernández

03/13/2017

PINS2017, SLAC

# Outline

- **The  $\nu$  Oscillation Phenomenon**
  - $\nu$  Energy Reconstruction
  - $\nu$  Interactions
- $\nu$  ‘anomalies’
- Precision era: LArTPC
- The SBN Program
  - The MicroBooNE Detector
  - MicroBooNE  $\nu$  Physics
  - MicroBooNE NuMI  $\nu$ ’s



# The $\nu$ Oscillation Phenomenon

Oscillation probability

$$P(\nu_\alpha \rightarrow \nu_\beta) = F(U_{PMNS}, \Delta m_{ij}^2, E_{\nu_\alpha}, L)$$

$$\Phi_{\nu_\alpha}^{produced}(E_{\nu_\alpha}) \otimes \sigma^{produced}(E_{\nu_\alpha}) \otimes \varepsilon^{produced}(E_{\nu_\alpha})$$



$$\Phi_{\nu_\beta}^{propagated}(E_{\nu_\beta}, \theta, \Delta m^2, \delta) \otimes \sigma^{propagated}(E_{\nu_\beta}) \otimes \varepsilon^{propagated}(E_{\nu_\beta})$$

$\nu_\alpha$

$\nu_\beta, \nu_\alpha$

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$\nu$  oscillations introduce differences in the flux spectrum and the ratio does not cancel the cross sections

$$\Phi_{\nu_\beta}^{propagated}(E_{\nu_\beta}, \theta, \Delta m^2, \delta) \otimes \sigma^{propagated}(E_{\nu_\beta}) \otimes \varepsilon^{propagated}(E_{\nu_\beta})$$

$$\frac{N_{events}^{propagated}(E_\nu)}{N_{events}^{produced}(E_\nu)} = \frac{\int \sigma(E'_\nu) \Phi(E'_\nu) P(E_\nu | E'_\nu) P_{osc}(E'_\nu) dE'_\nu}{\int \sigma(E'_\nu) \Phi(E'_\nu) P(E_\nu | E'_\nu) dE'_\nu}$$



# The $\nu$ Oscillation Phenomenon

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### Main limitations:

- $\nu$ -Nucleus Interaction Cross Section
- Flux Uncertainties

### One detector scenario

In general, the detector doesn't cover all the phase-space of generated neutrinos, and we may have un-predicted migrations along neutrino spectra.

Even if full phase-space is covered, cross section distributions for a non-oscillating hypothesis are still not fully understood.

# The $\nu$ Oscillation Phenomenon

## Oscillation probability

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### Main limitations:

- $\nu$ -Nucleus Interaction Cross Section
- Flux Uncertainties

With 2 detectors, the near one can be used to understand neutrino-nucleus interaction effects. However, even if the 2 detectors use the same target, since neutrino energy spectra is not the same, we cannot cancel cross section uncertainties.

Two/more detector scenario



# $\nu$ Energy Reconstruction

Since we don't have a monochromatic  $\nu$  beam we deal with the  $\nu$  energy reconstruction:

- $E_\nu$  relies on lepton kinematics
  - Different methods, relying on the reaction type description or final state particles
- Either if we use an exclusive or inclusive channel in the measurement,
  - hadron composition and its kinematics are strongly affected by FSI
- Fermi momentum, Pauli blocking and bound energy plays significant role
- Neutrons are not typically accessible by current  $\nu$  detectors.

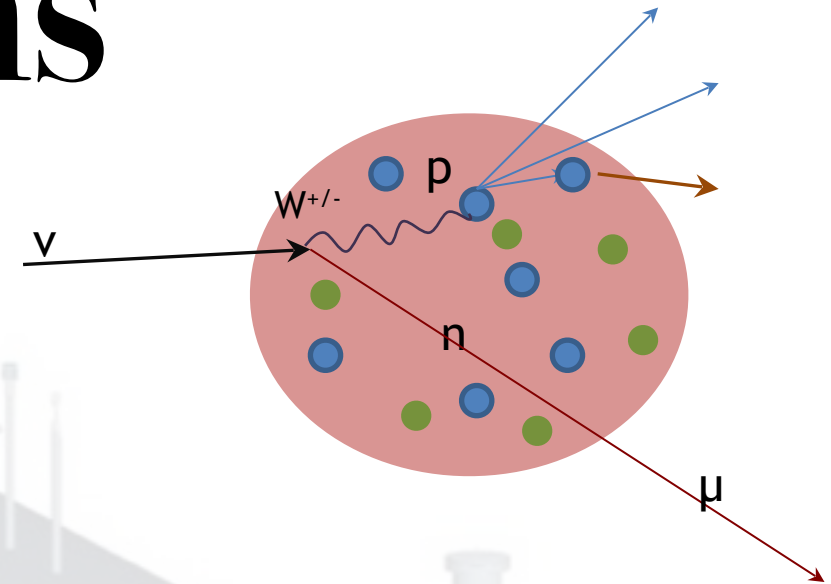
Some experiments (i.e.  $\nu_e$  appearance @T2K) use only lepton kinematics in the oscillation fit to reduce un-knows.

But still, even if this reduces the extrapolation, lepton kinematics are not yet fully understood.

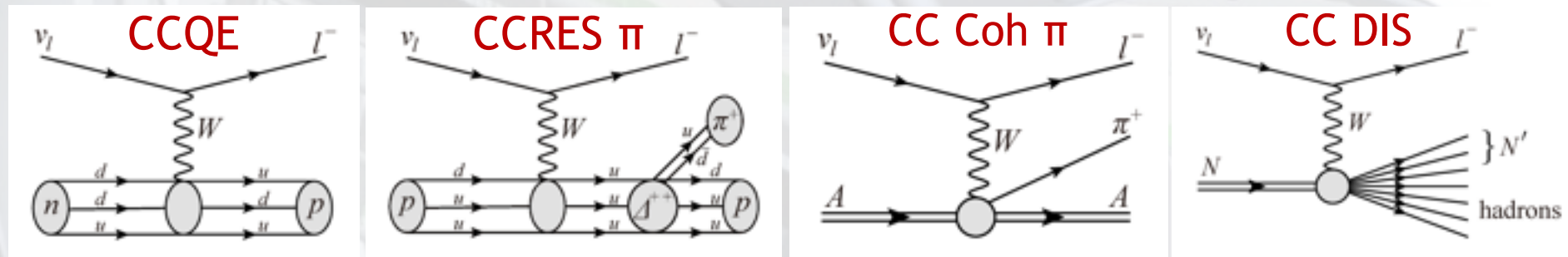
Because life is not easy...

# $\nu$ Interactions

To characterize the un-oscillated neutrinos  
we use CC events



Interaction neutrino-nucleon



Interaction with the nucleus

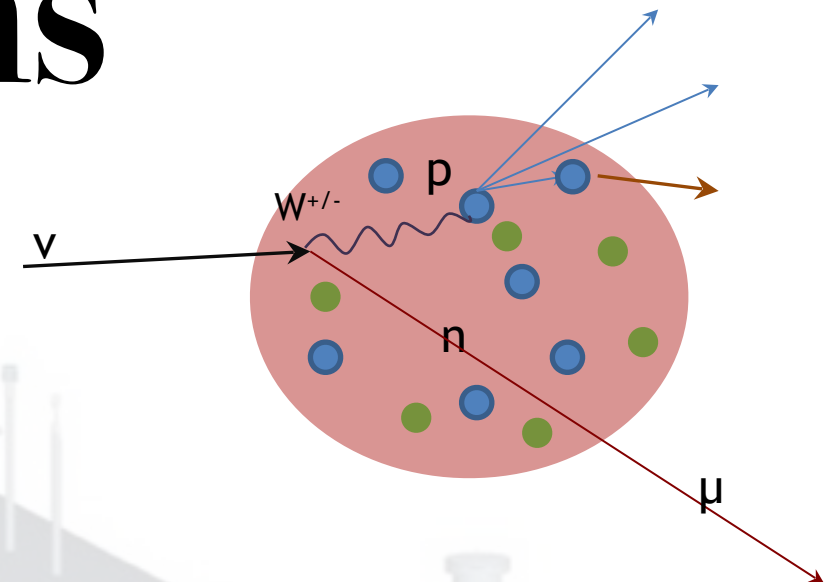
- **Final State Interaction:**
  - Pion absorption
  - Proton, pion scattering
  - **Charge ex-change**
- **Nuclear Short-Range Effects**
  - **Meson Ex-change Current**
- **Nuclear Long-Range Effects**
- **Nuclear Modelling**
  - Fermi motion
  - Binding energy
  - Pauli blocking



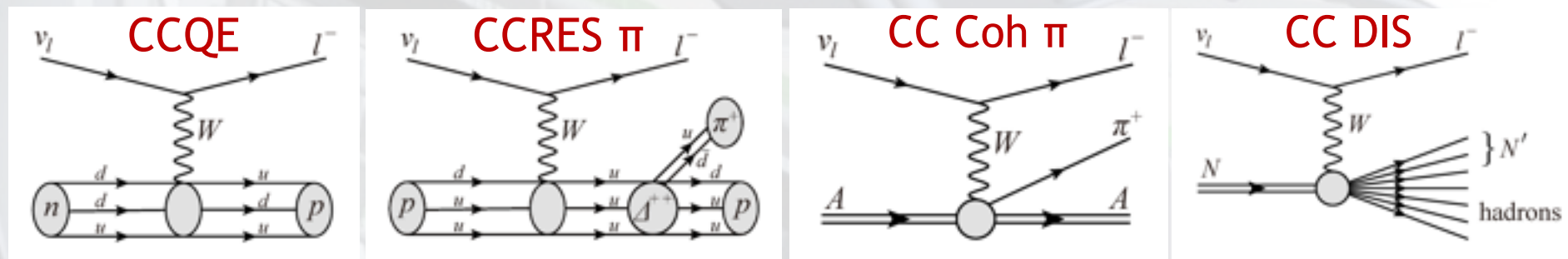
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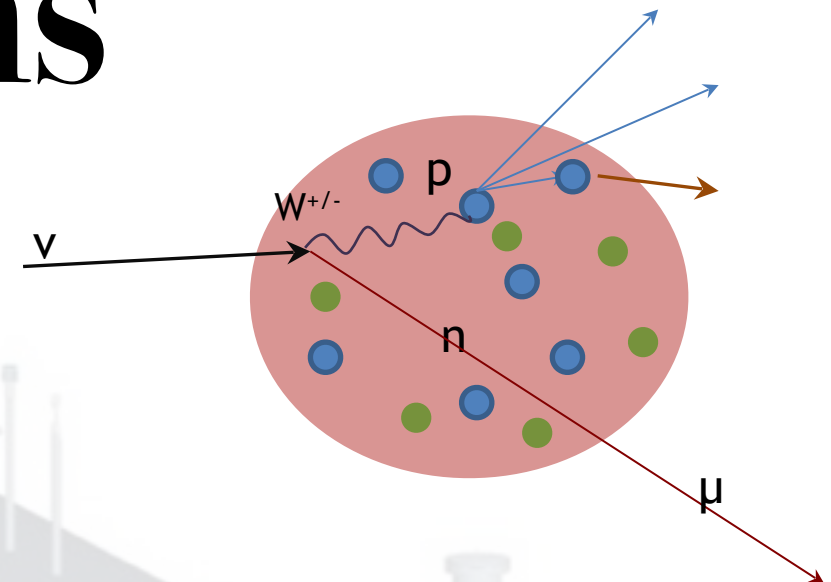
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- $\nu_e/\nu_\mu$

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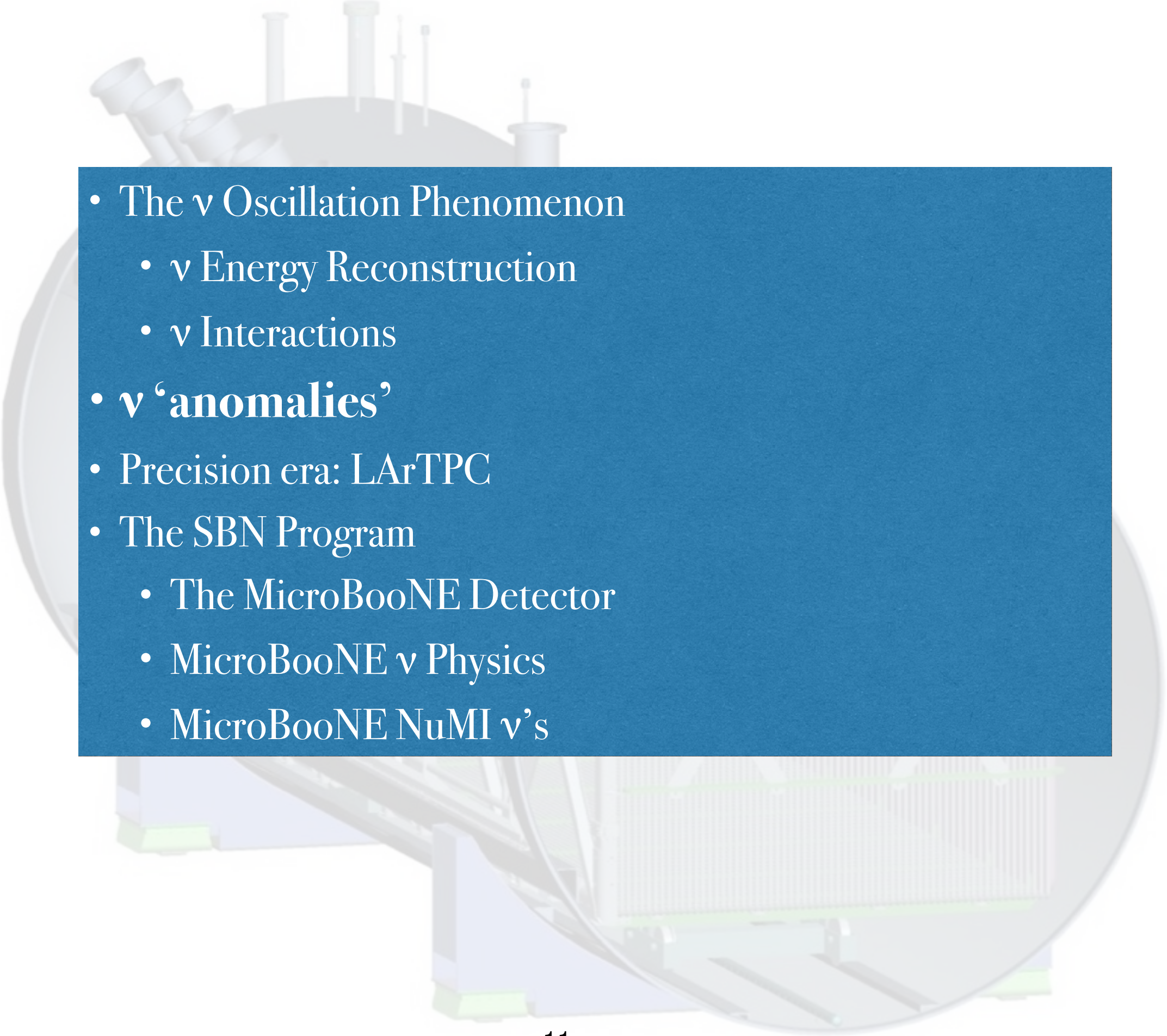


Known limitations for neutrino-carbon/water experiments (T2K, NOvA, MiniBooNE) in osc. analysis.

For argon we expect most of these effects to be bigger

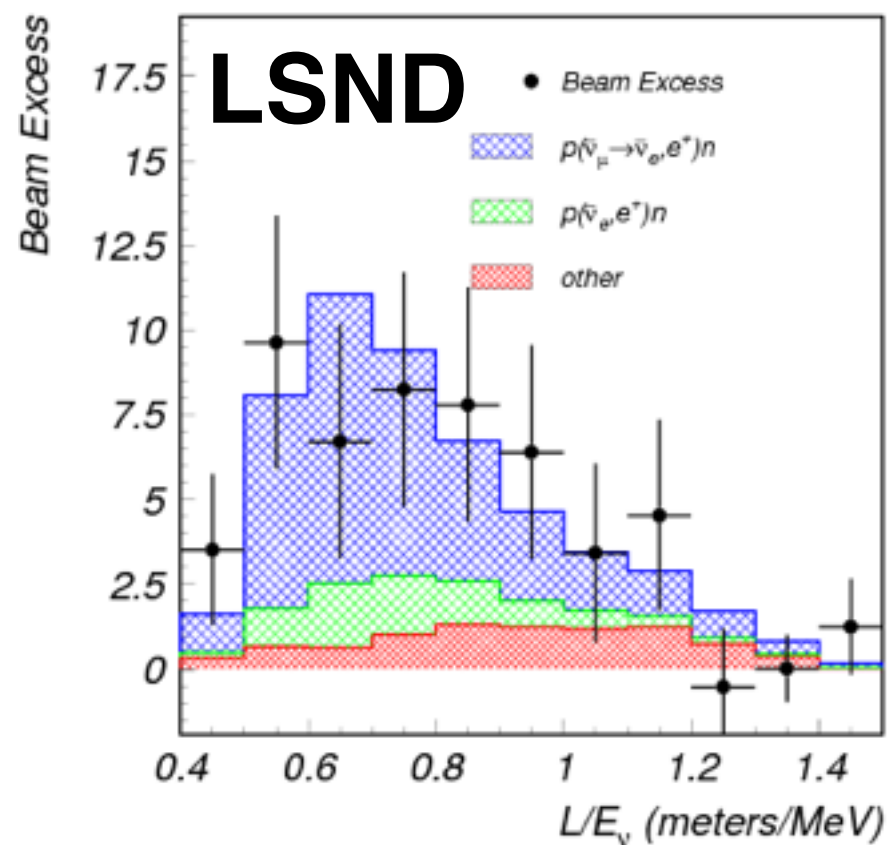
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- 
- The  $\nu$  Oscillation Phenomenon
    - $\nu$  Energy Reconstruction
    - $\nu$  Interactions
  - **$\nu$  ‘anomalies’**
  - Precision era: LArTPC
  - The SBN Program
    - The MicroBooNE Detector
    - MicroBooNE  $\nu$  Physics
    - MicroBooNE NuMI  $\nu$ ’s

# $\nu$ 'anomalies'

In addition to all... do we have only 3 neutrino families?  
Within the Standard Model we have measured 3 families.  
...and in the meanwhile, experiments observe unpredicted events...



Phys.Rev. D64 (2001) 112007

$$\bar{\nu}_\mu \rightarrow \bar{\nu}_e?$$

**Non-interacting neutrinos? (i.e. sterile)**

Need to see similar effect in  
neutrino mode to discard CP  
and  $\nu_\mu$  to discard lepton flavor  
differences.

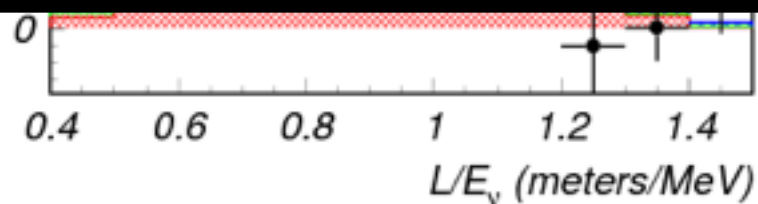


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Within the Standard Model we have measured 3 families.  
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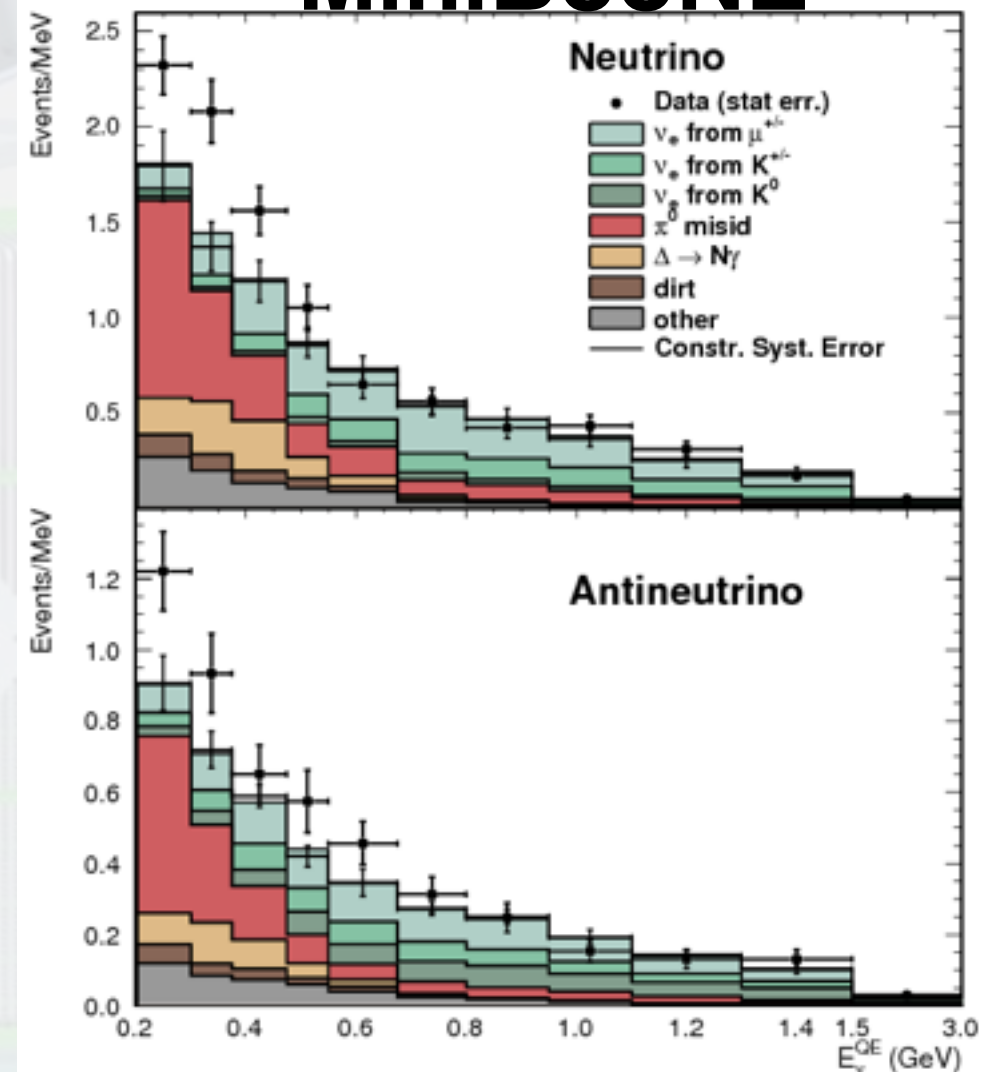
To identify origin of the low energy events  
we need higher phase-space coverage & reduced background:

- Identify  $e^-/\gamma$ 
  - Crucial to reduce background
- Lower energetic particles should be accessible



Phys.Rev. D64 (2001) 112007

## MiniBooNE



Phys.Rev.Lett. 110 (2013) 161801

# $\nu$ 'anomalies'

Up to now, MiniBooNE data/MC discrepancies:

**Pion production uncertainties**  
 $\nu$  induced  $\pi^0$  production underestimations,  
observed in other experiments (e.g. MINERvA)

**More  $\nu$ -nucleus scattering un-knows?**

**More exotic physics (i.e. heavy  $\nu$ )**

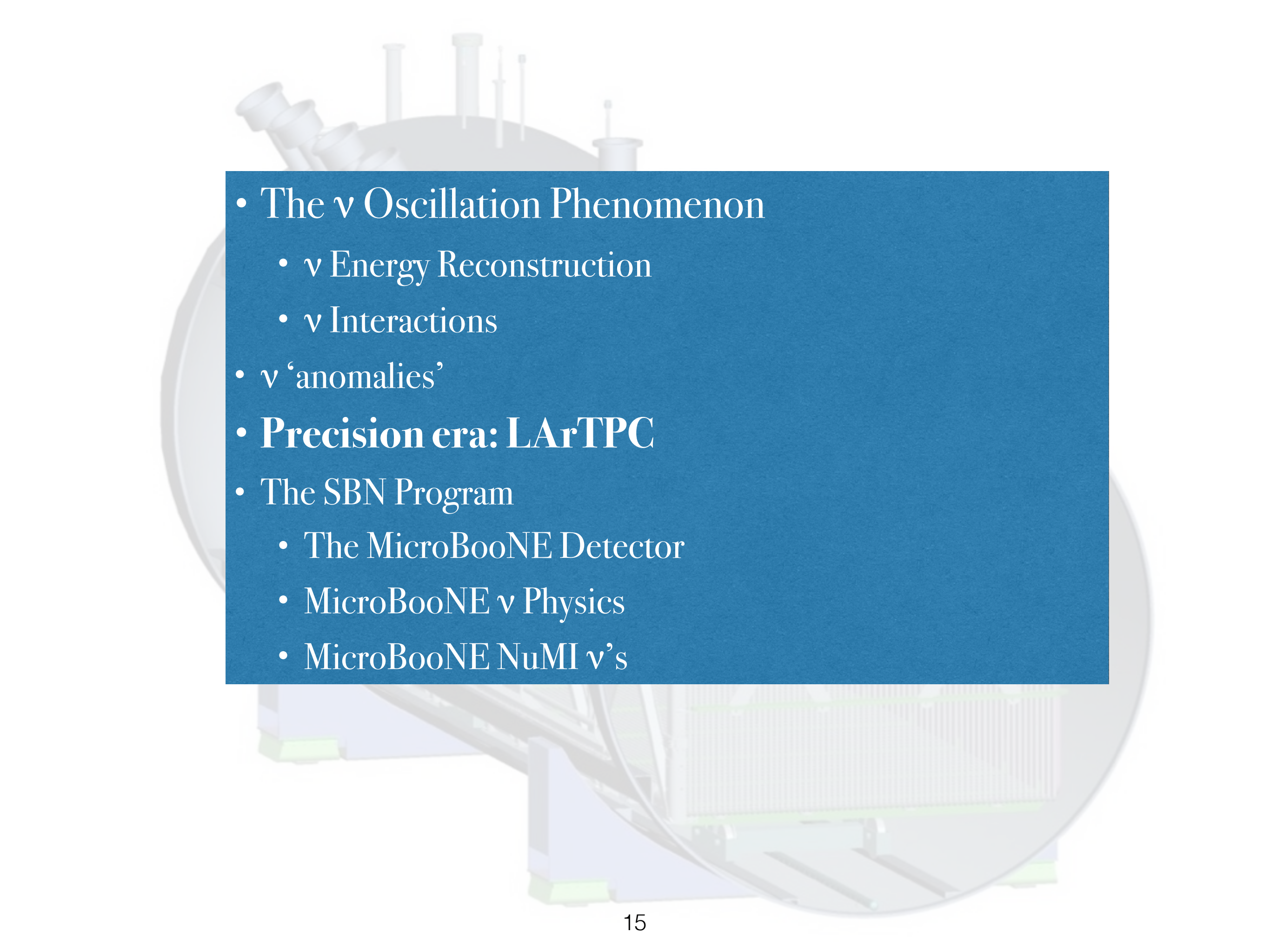
**Sterile neutrino(s): 3+1, 3+2, 3+3 scenarios**

**Clear identification of  $e^-/{}^+/\gamma$  events**

**Understand differences  $\nu_e$ - $\nu_\mu$**

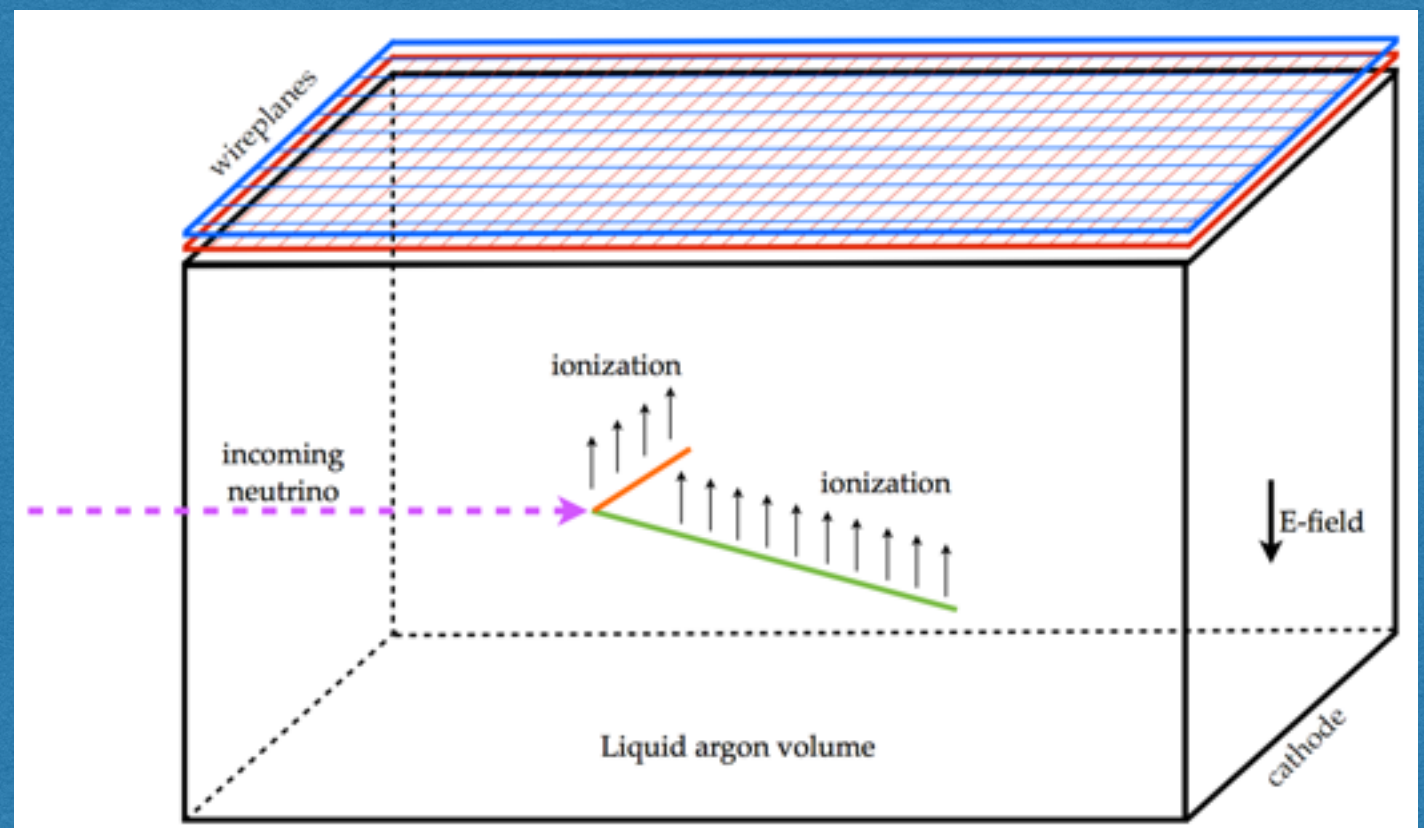
**Better model constraints in  $\nu$ -nucleus  
scattering**



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  - **Precision era: LArTPC**
  - The SBN Program
    - The MicroBooNE Detector
    - MicroBooNE  $\nu$  Physics
    - MicroBooNE NuMI  $\nu$ ’s

# Precision era: Liquid Argon TPC

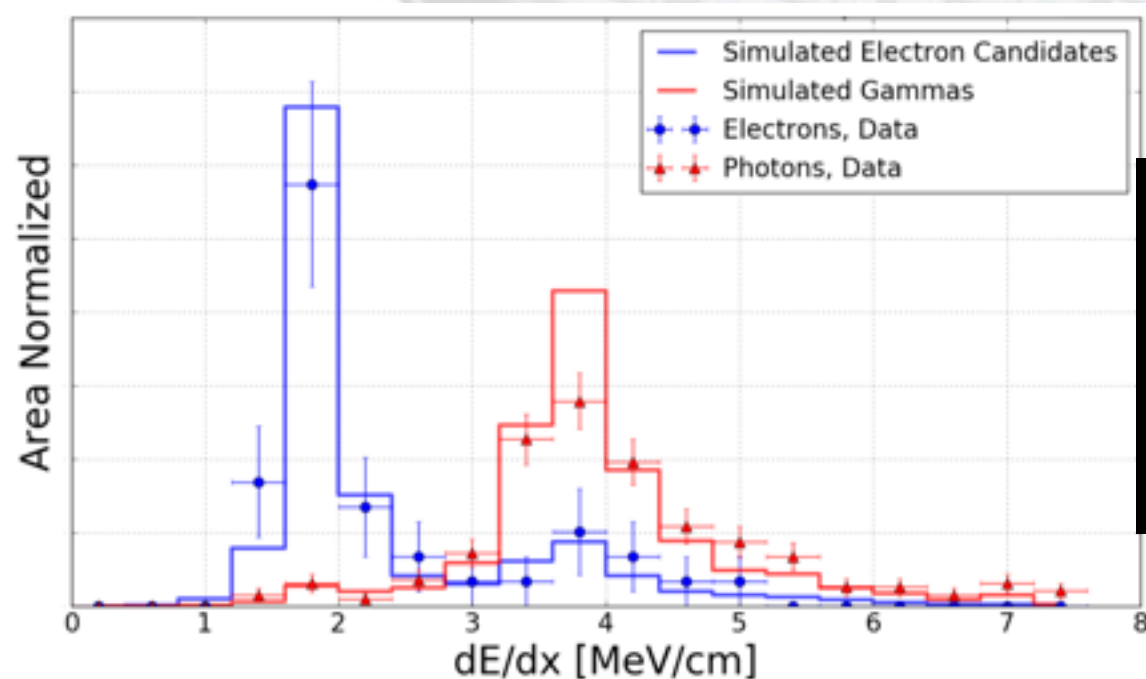
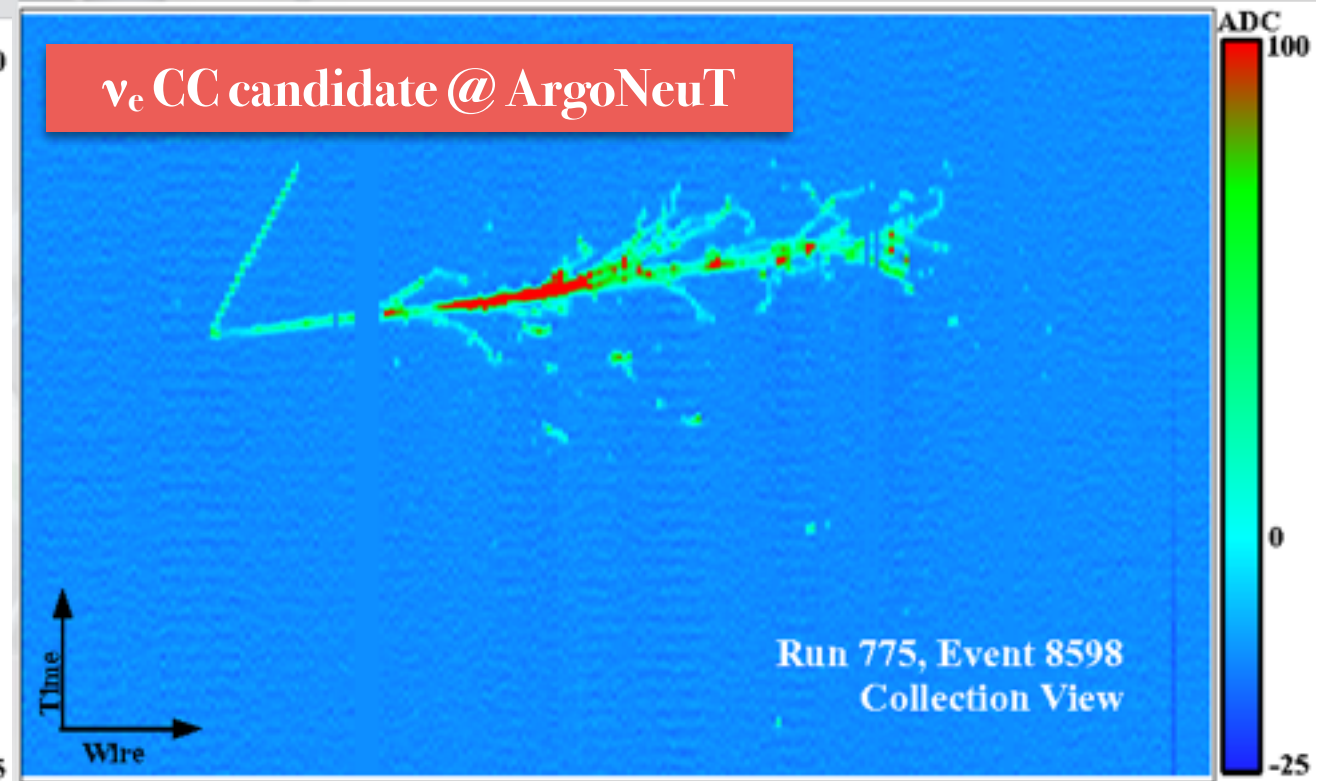
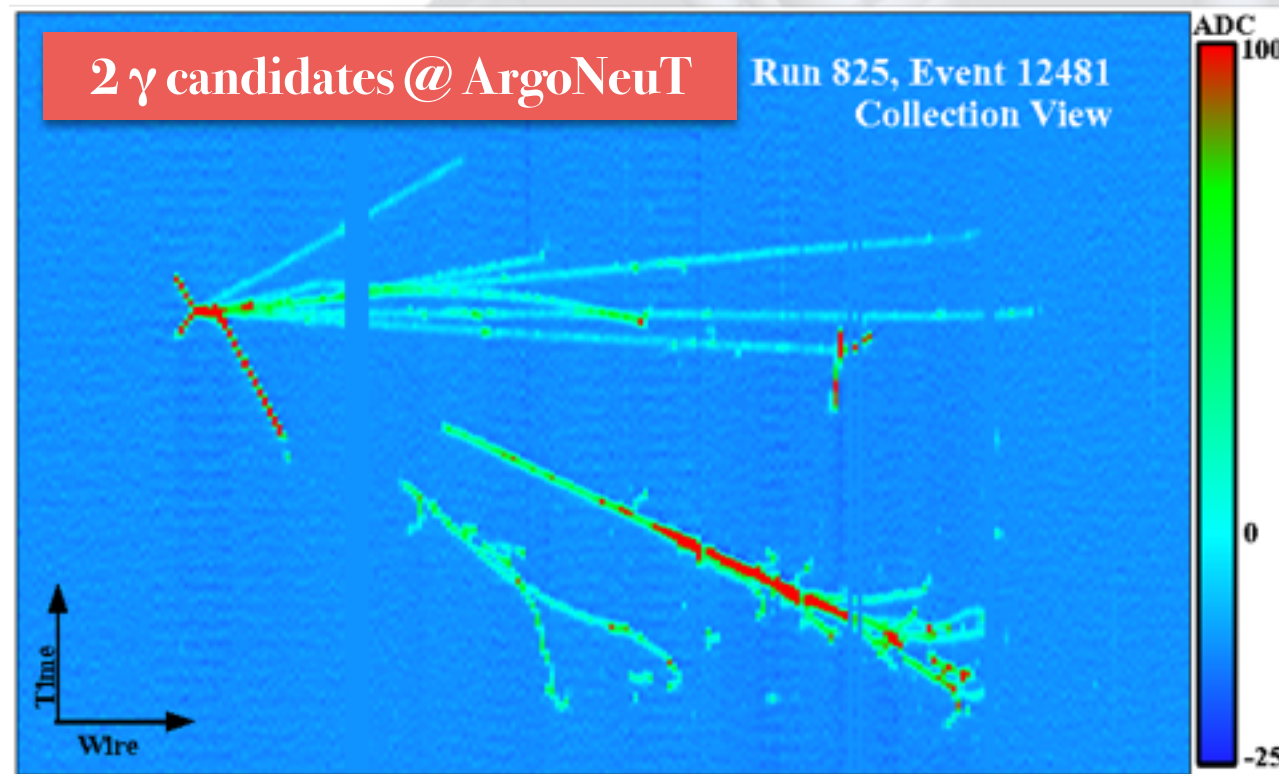
- Ionization from traversing charged particles is drifted along E-field to the segmented wire planes.
  - argon ionizes easily,  $\sim 70 \text{ ke/cm}$  (@500 V/cm)
- Wire pulse timing information is combined with known drift speed to determine drift-direction coordinate.
- Calorimetry information is extracted from wire pulse characteristics.
- Abundant **scintillation light**, which LAr is transparent to, also available for collection and triggering.
  - $40 \text{ k } \gamma/\text{MeV}$  @null E-field
- Argon is **40%** more dense than water.
- 1% abundance in the atmosphere.



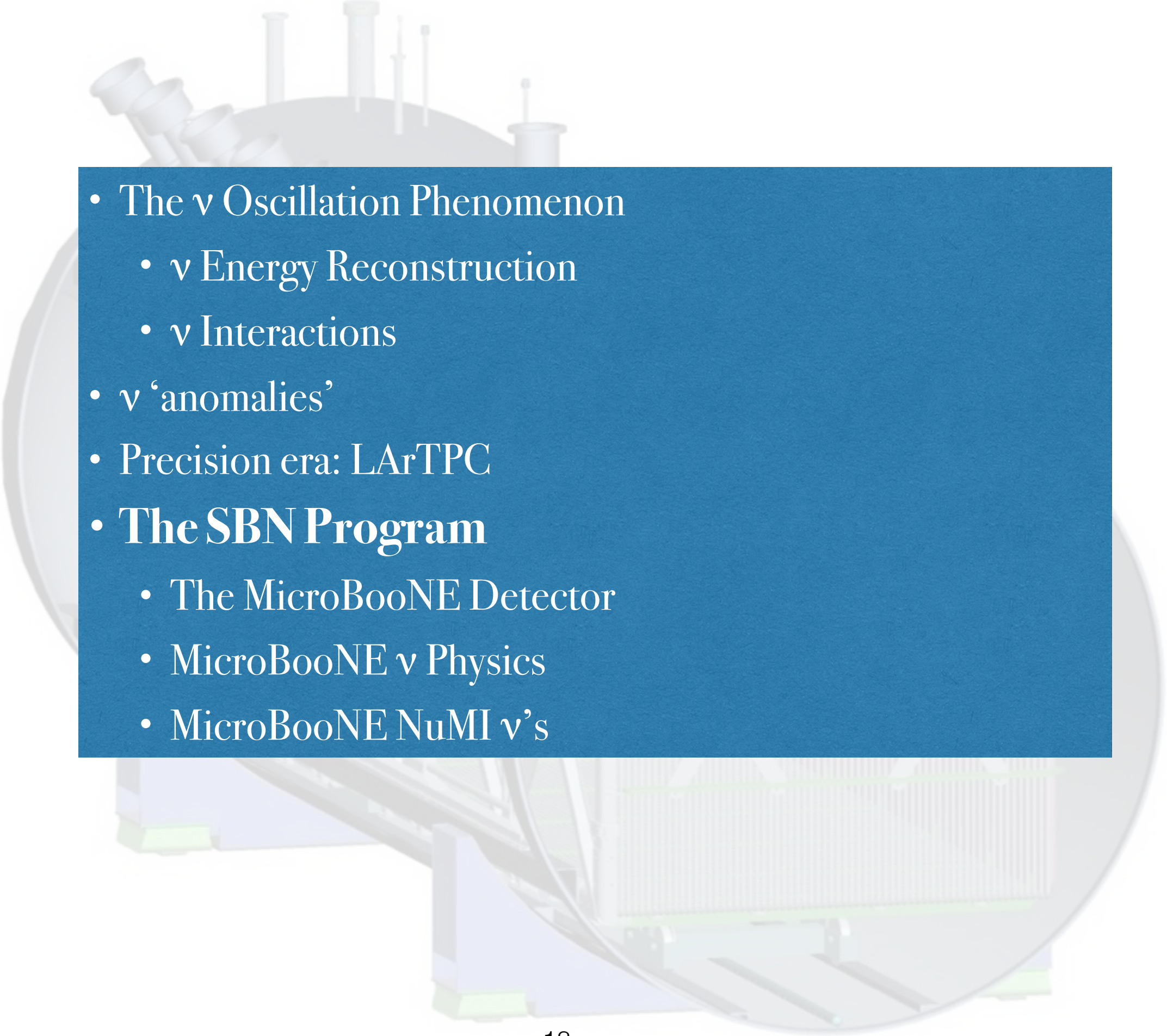


# Precision era: Liquid Argon TPC

Phys.Rev.D arXiv:1610.04102



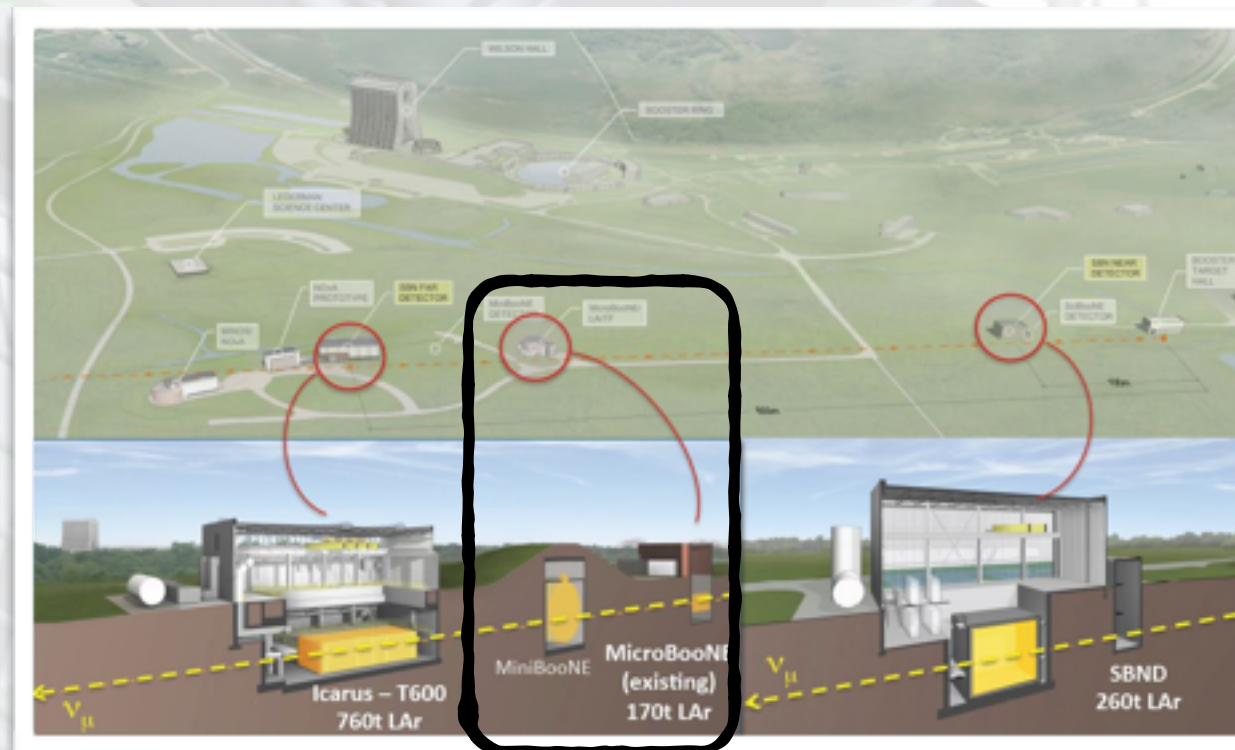
Ability to identify particles at any  
direction from @20MeV  
Distinguishable signature  $e^-/\gamma$

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    - The MicroBooNE Detector
    - MicroBooNE  $\nu$  Physics
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# The SBN Program

SBN (Short Baseline Neutrino) aims to search for non-standard  $\nu$  oscillations by  $\nu_e$  appearance and  $\nu_\mu$  disappearance with unprecedented precision in BNB.



## Main MicroBooNE physics goals:

- Investigate MiniBooNE low-energy un-predicted data ( $\nu_e$  CC events)
- Measure first low-energy  $\nu$ -Ar cross sections
- R&D for Deep Underground Neutrino Experiment (DUNE)
- Joint oscillation analysis within the Short Baseline Neutrino (SBN) program

+ Exotic physics capability studies (proton decay, SN,...). See Yun-Tse's talk this Thursday!

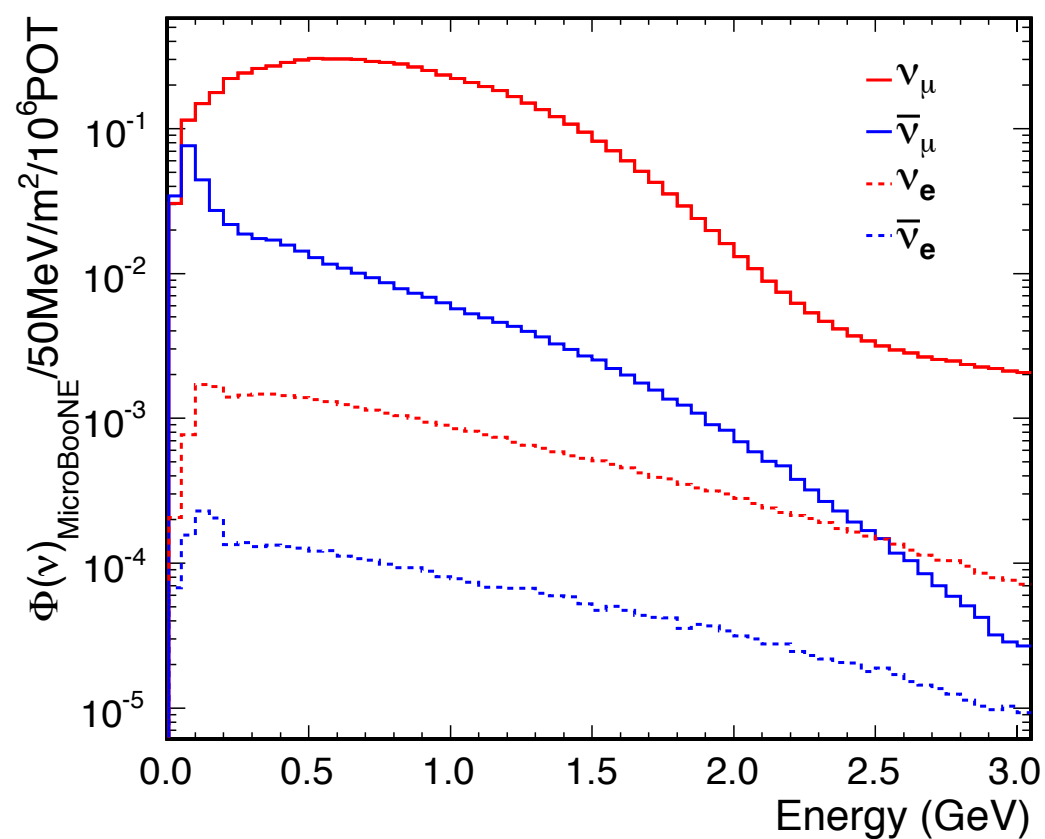
# The BNB $\nu$ beam

8 GeV Protons from BNB

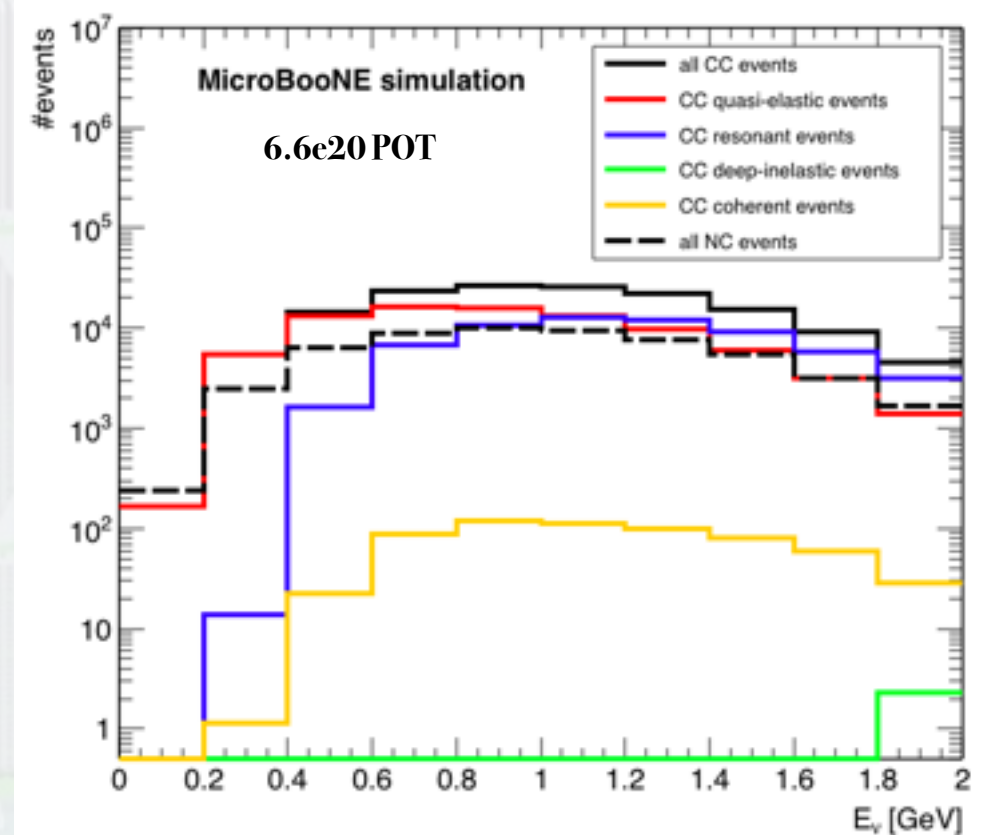
Almost pure  $\nu_\mu$  beam (0.5% intrinsic  $\nu_e$  contamination)

Located on-axis from  $\nu$  source @470m

Booster Neutrino Beam (BNB) flux at MicroBooNE



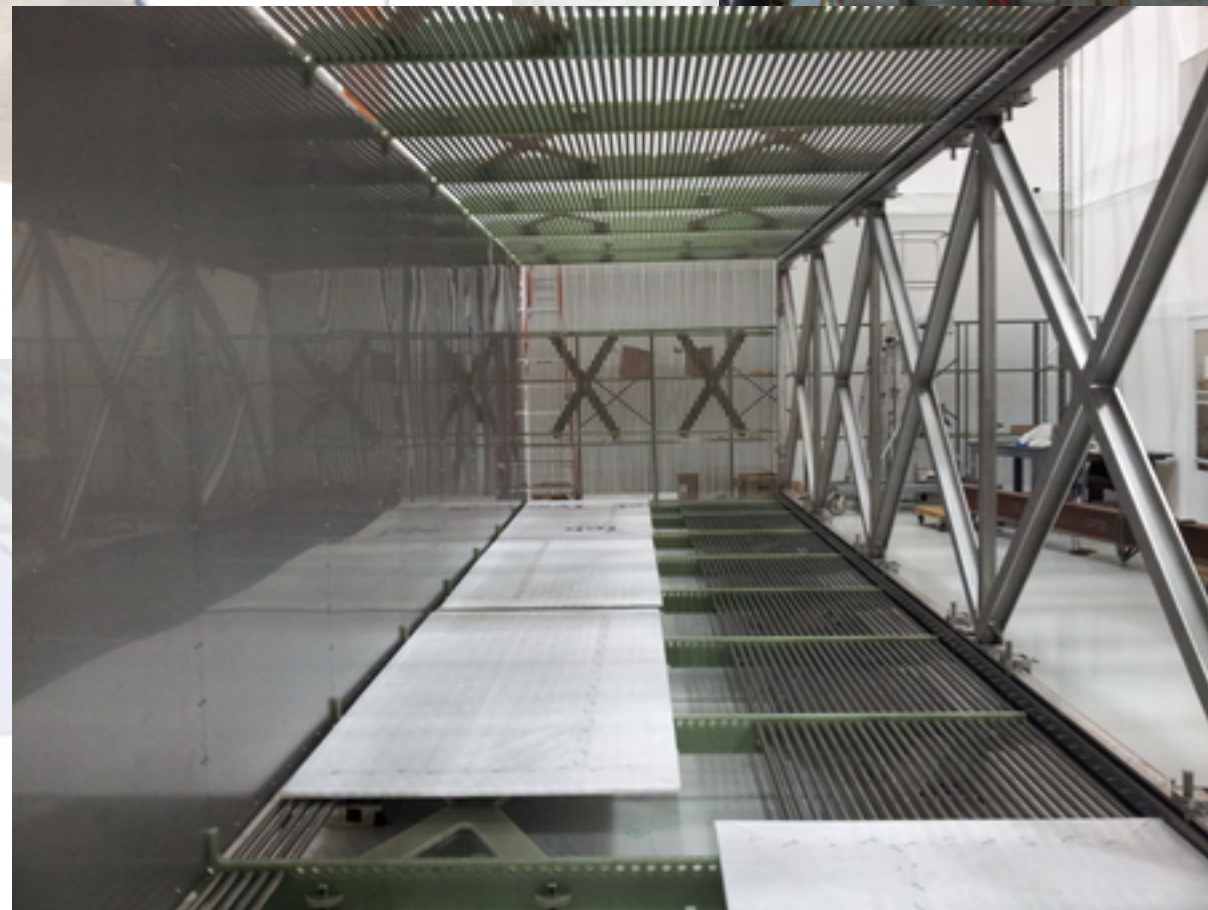
BNB  $\nu_\mu$  interactions in MicroBooNE





# The MicroBooNE Detector

## MicroBooster Neutrino Experiment



156 collaborators from 28 institutions (7 non-U.S.)

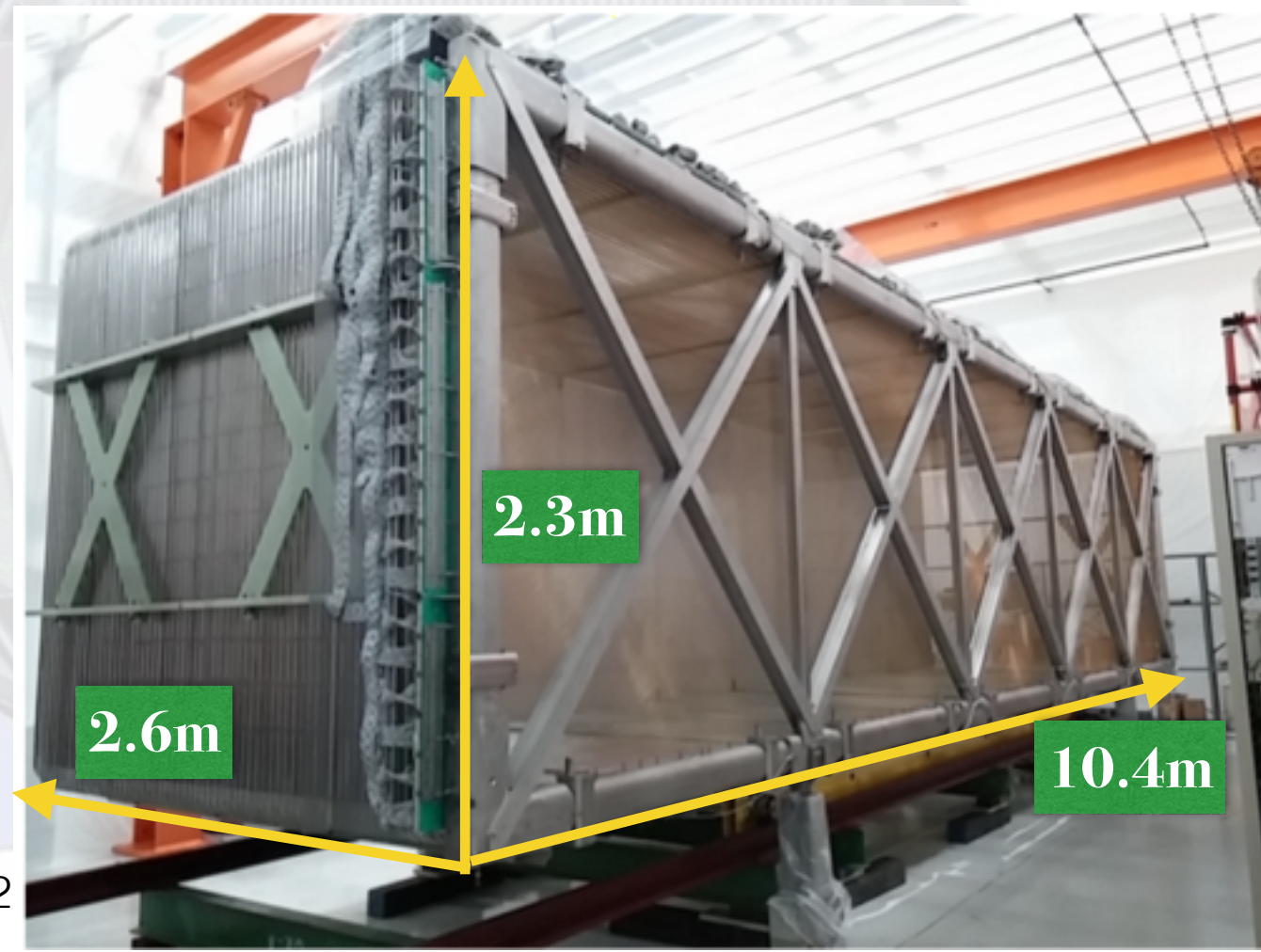


# The MicroBooNE Detector

- 170 tons of liquid argon
  - 86 tons of active mass
- Non-evacuated liquid argon fill
- Cold (in LAr) front-end electronics
- Near-surface operation
- UV laser calibration system

‘Design and Construction of the MicroBooNE Detector’

arXiv:1612.05824

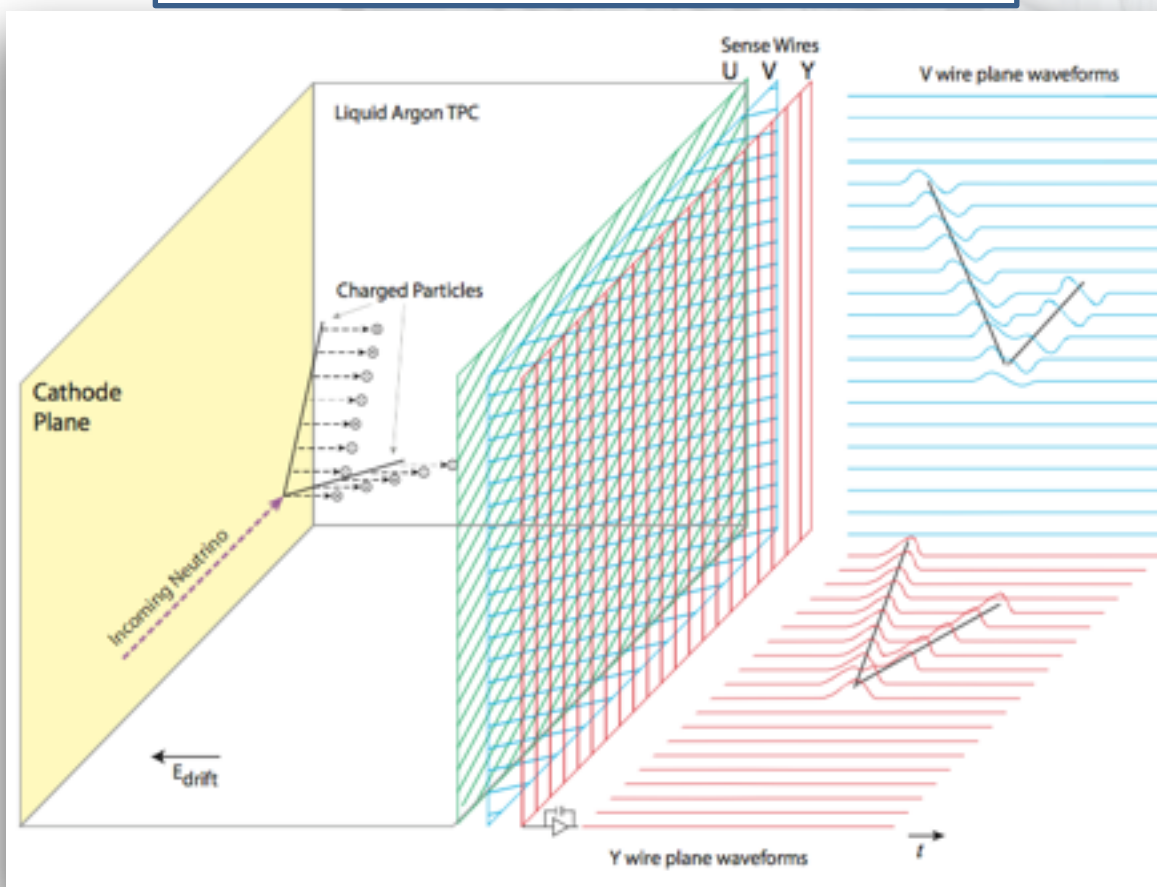




# MicroBooNE TPC

- 8192 wires, 3mm separation between wires
- Two induction planes (U, V) and one collection plane (Y), 3 mm pitch
  - signal on all three planes
  - Drift E-field at 273 V/cm, surrounding field cage allows E-field uniformity
  - 8000+ channels with front-end electronics in LAr
- 3D event reconstruction by combining signals from all 3 planes (2 required)

arXiv:1609.06169 [physics.ins-det]



**First large scale LArTPC to employ cold front end electronics**

- Suppresses noise

**RMS noise on the readout wires drops as the cryostat is cooled**

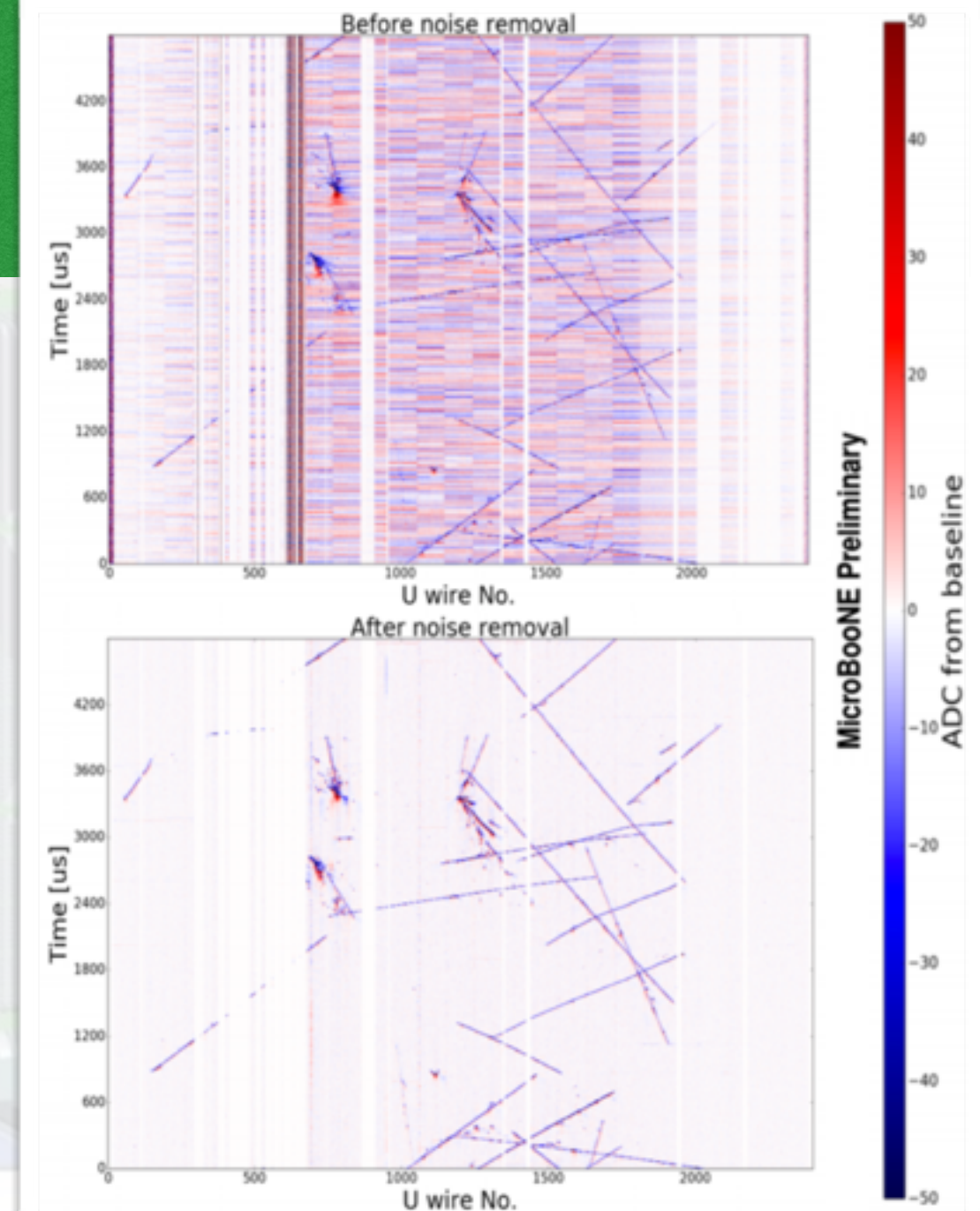
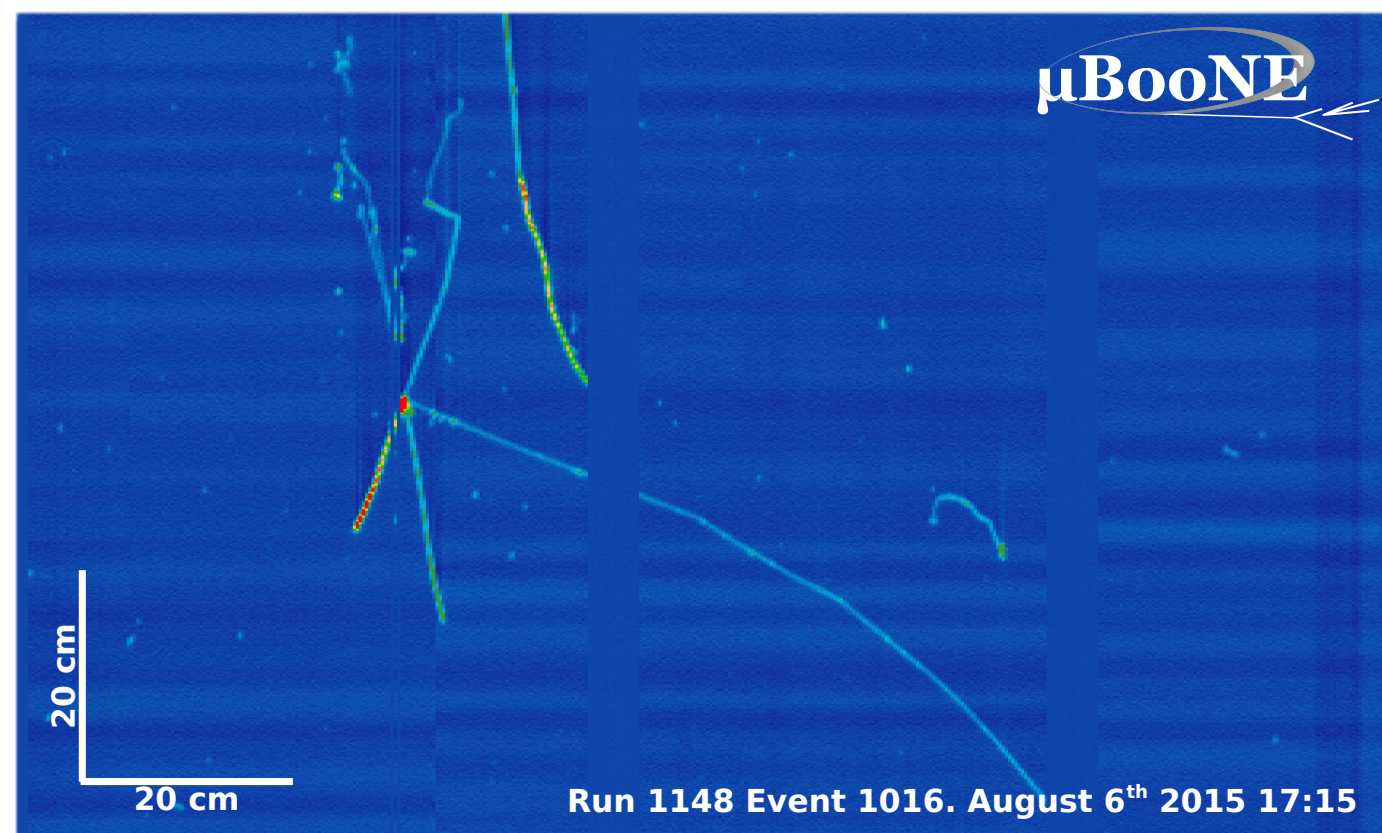
- The noise begins to raise once LAr is added due to change in capacitance



# MicroBooNE TPC

MICROBOONE-NOTE-1016-PUB

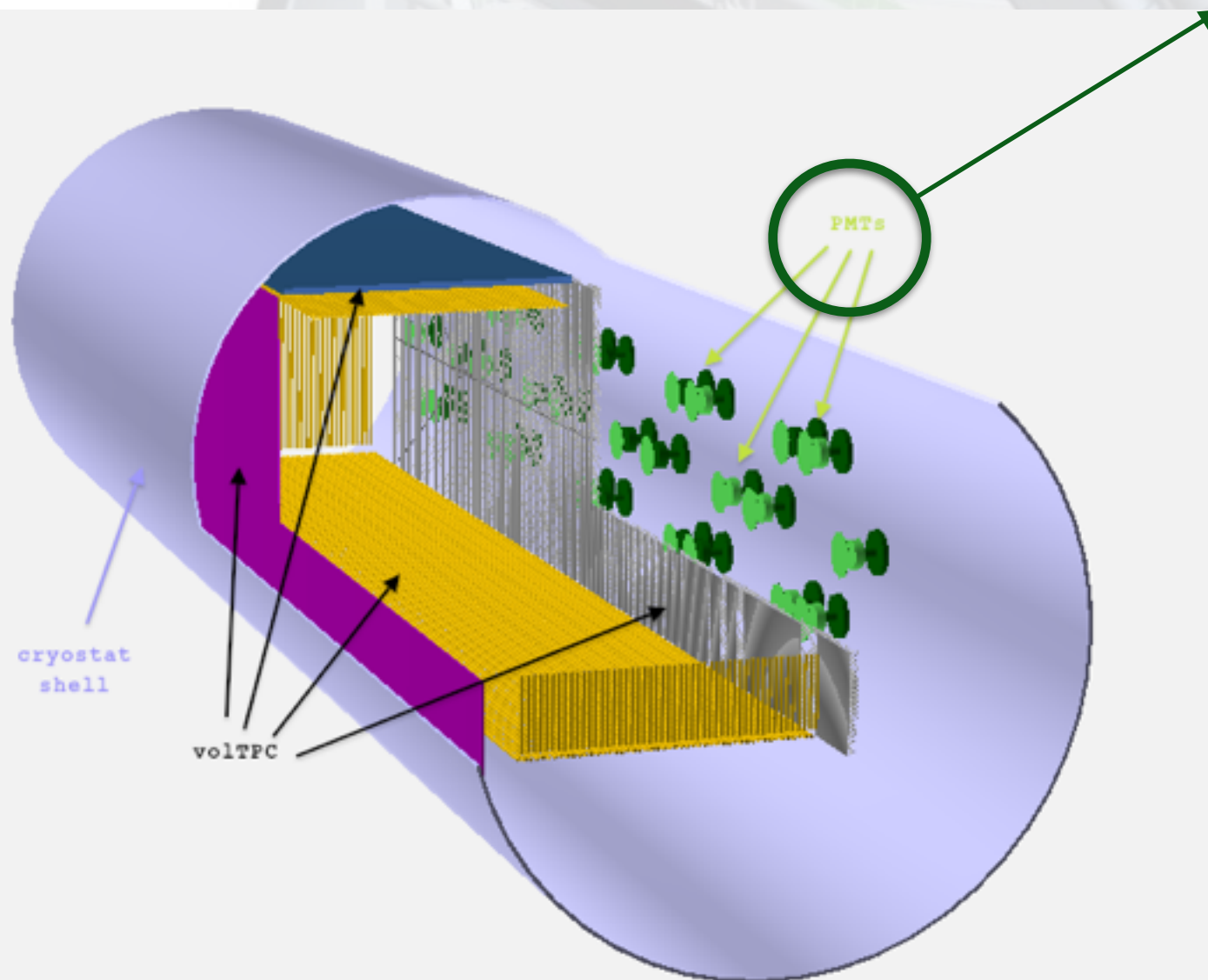
- Coherent noise over group of channels
  - This noise is associated to a voltage regulator on a warm service board
  - With software filtering we are able to improve signal-to-noise by factor of 2
- Signal-to-noise ratio after software noise filtering
- U plane 15.8 : 1
- V plane 12.9 : 1
- Y plane 45.3 : 1



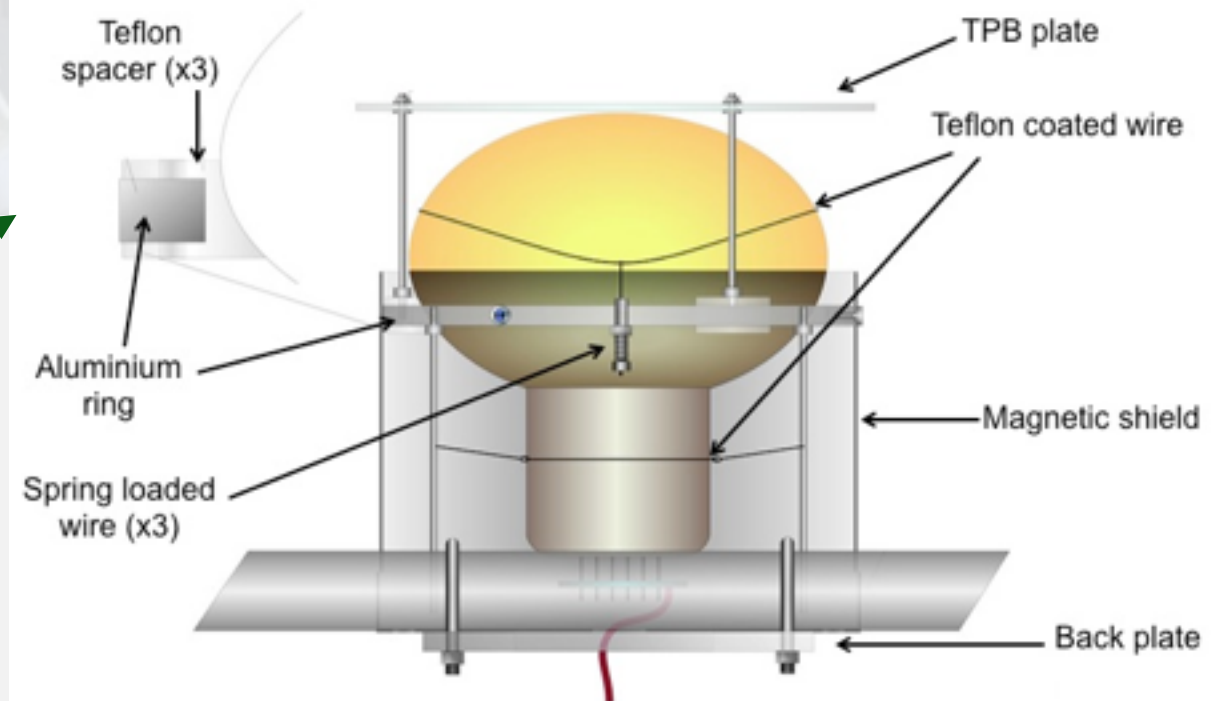


# MicroBooNE Photon Detectors

32 PMTs + 4 light guide paddles



PMT mount with a PMT



A wire ring is pulled down by 3 spring loaded wires to an aluminium ring.  
Direct contact of the PMT to the aluminium ring is avoided by Teflon blocks.  
The magnetic shield and the TPB plate are fixed on to the PMT mount.



# MicroBooNE Physics

$\mu$ BooNE

Finished commissioning August 2015  
Taking  $\nu$  data since October 2015

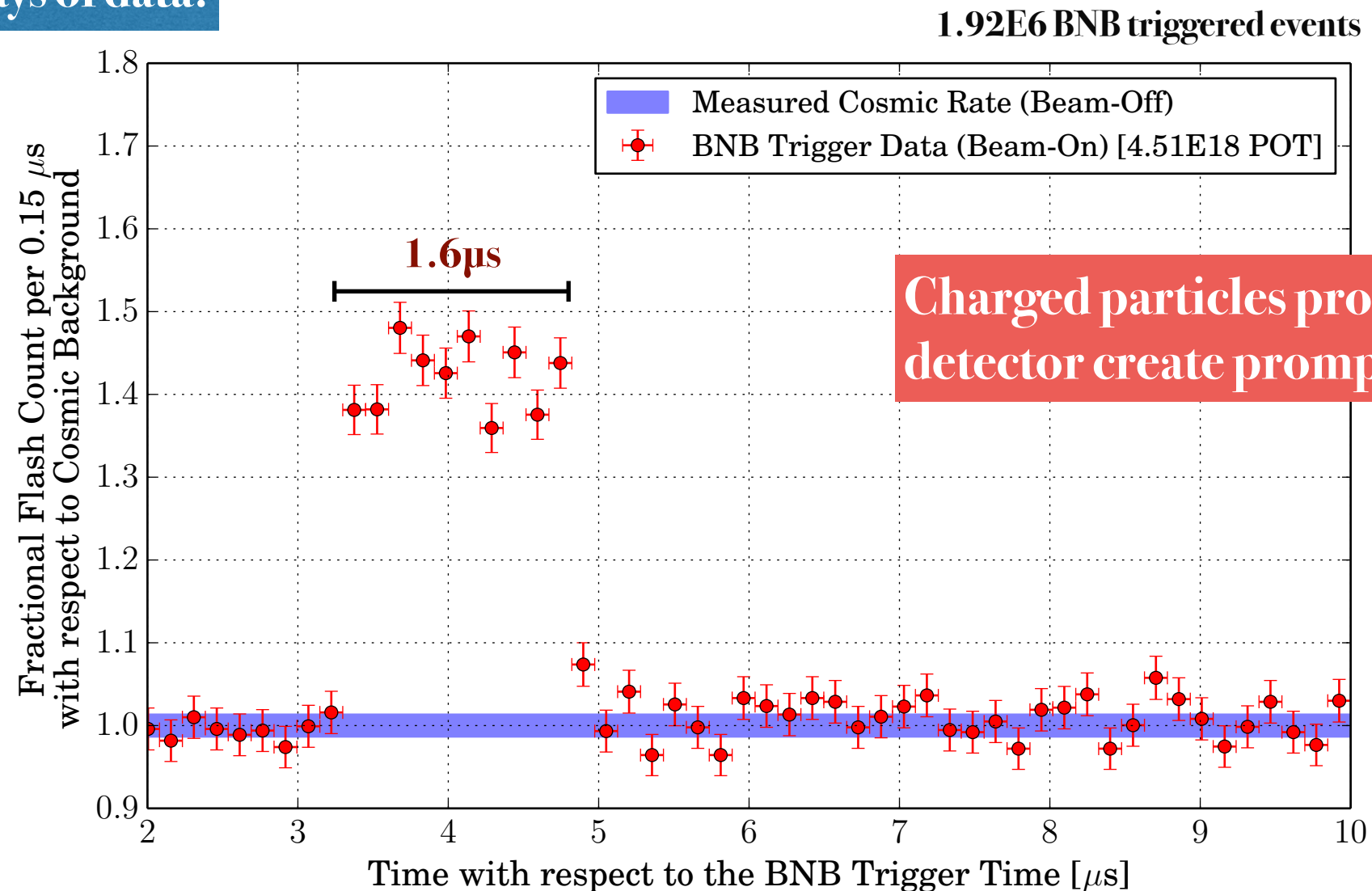
75 cm

Run 3493 Event 41075, October 23<sup>rd</sup>, 2015



# MicroBooNE Physics

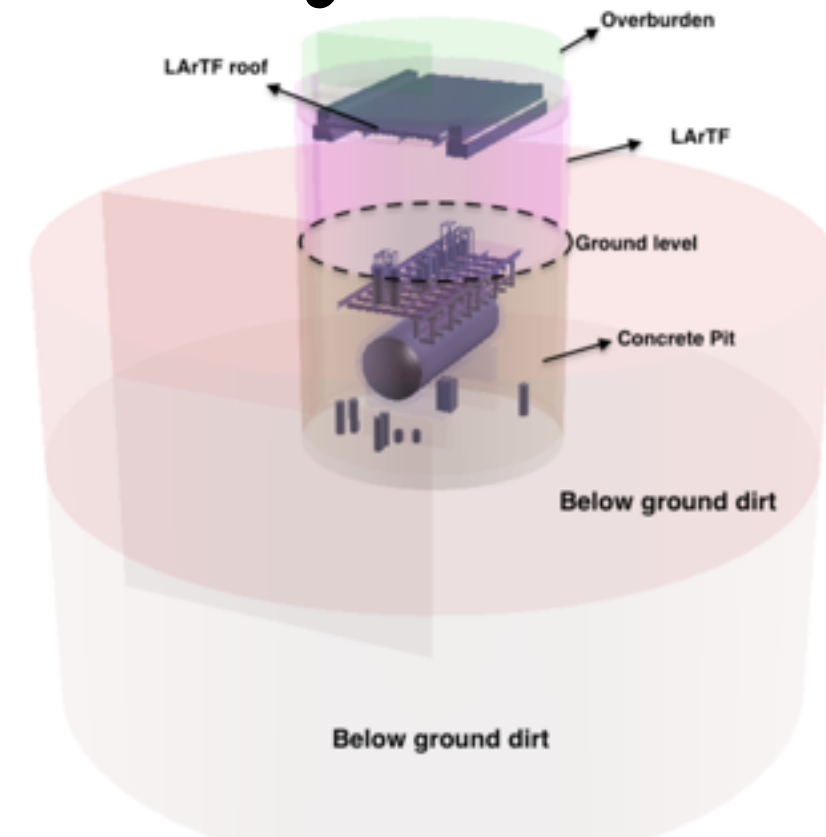
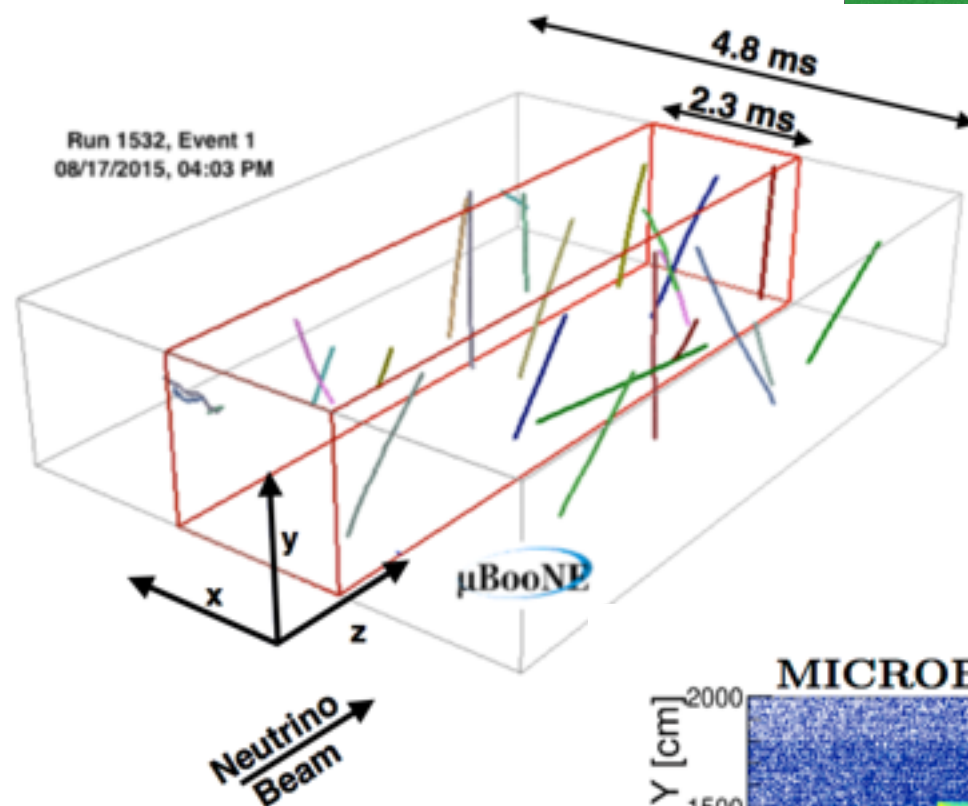
From the first days of data!



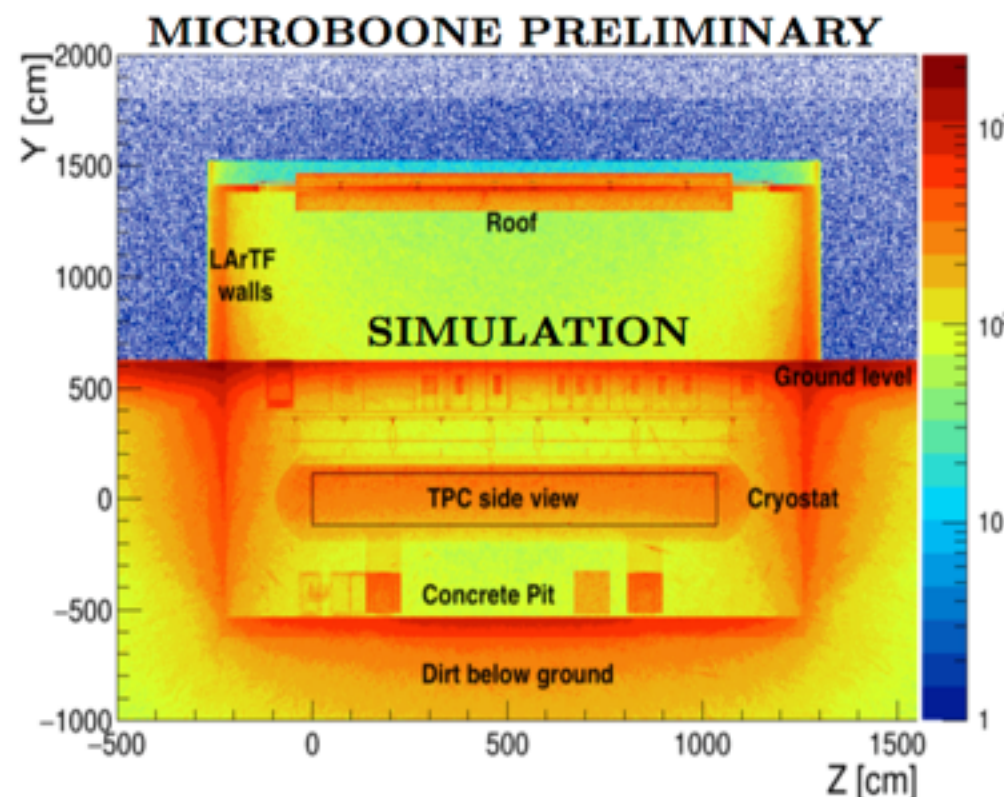
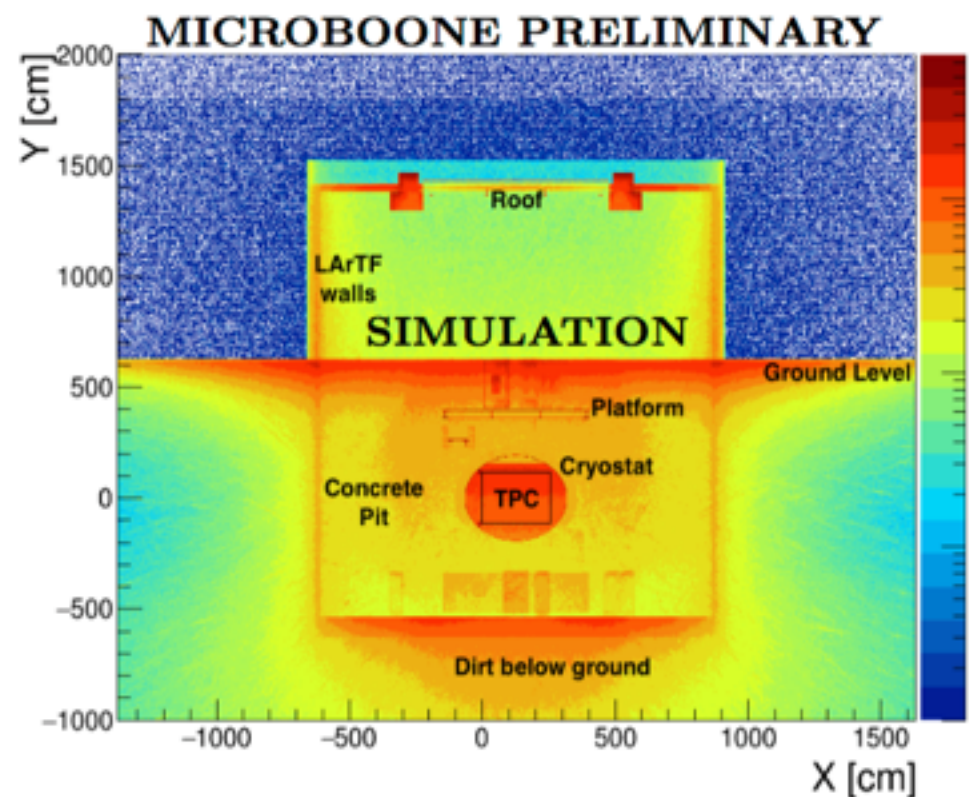
A clear excess can be seen due to neutrinos between 3 and 5  $\mu\text{s}$  after the trigger.

# MicroBooNE Cosmic Physics

Cosmic data is used for the characterization of the detector.



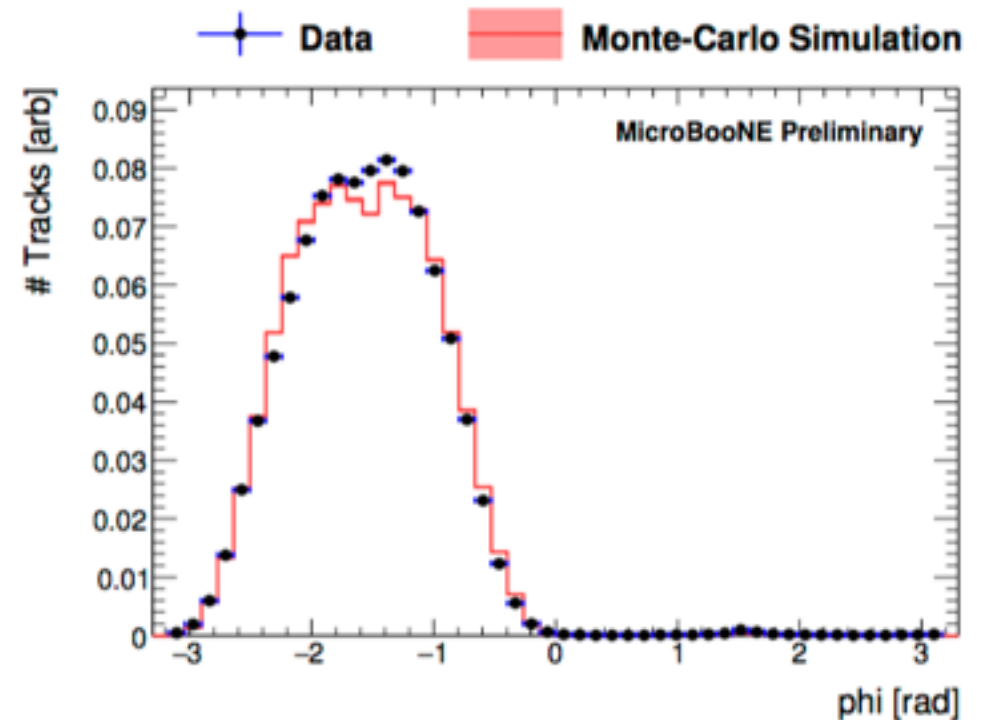
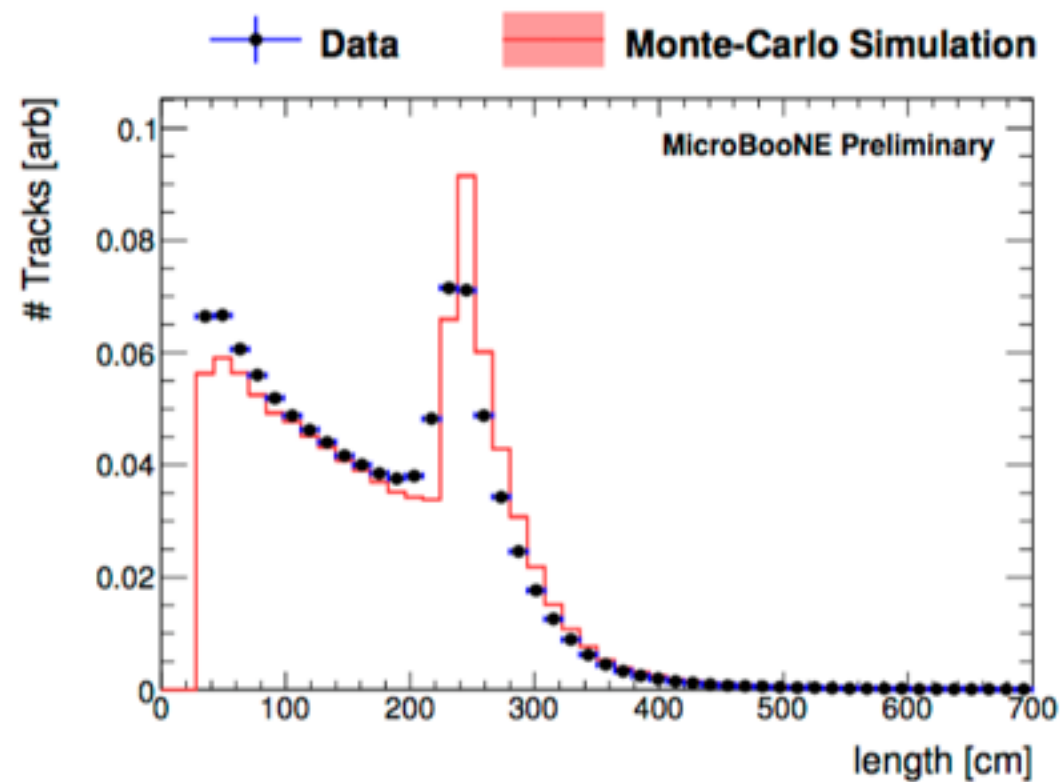
drift time of 2.3 ms  
readout time 4.8 ms





# MicroBooNE Cosmic Physics

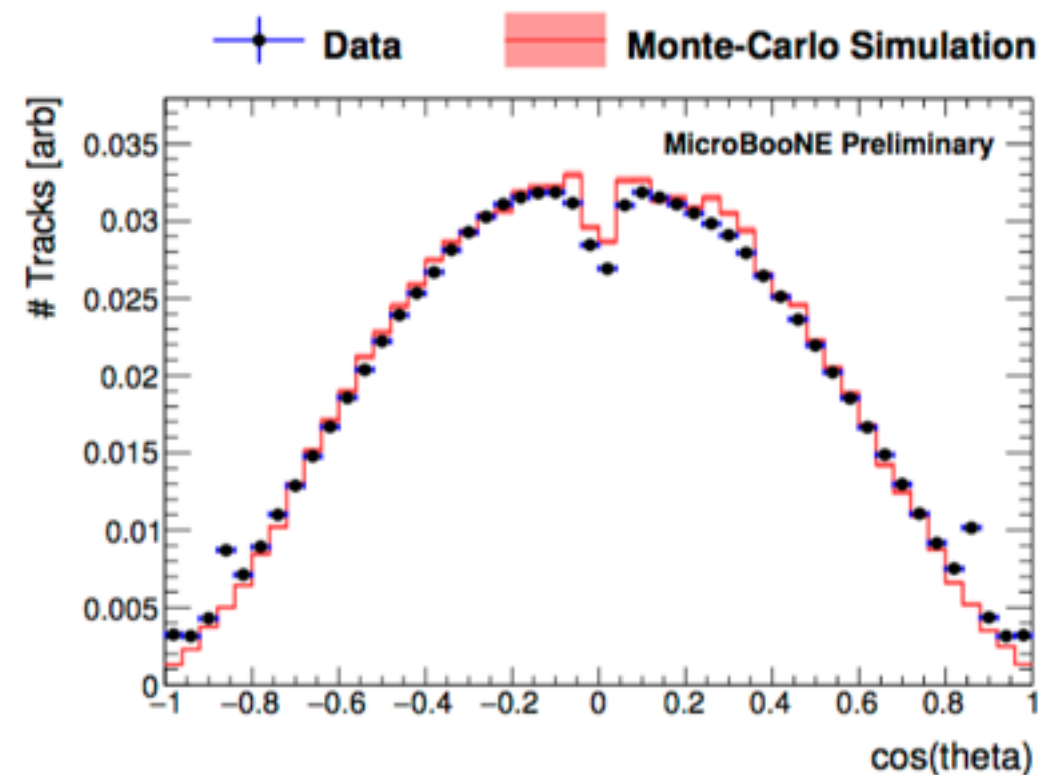
We combine hit information from the 3 planes to reconstruct 3D objects



Area normalized

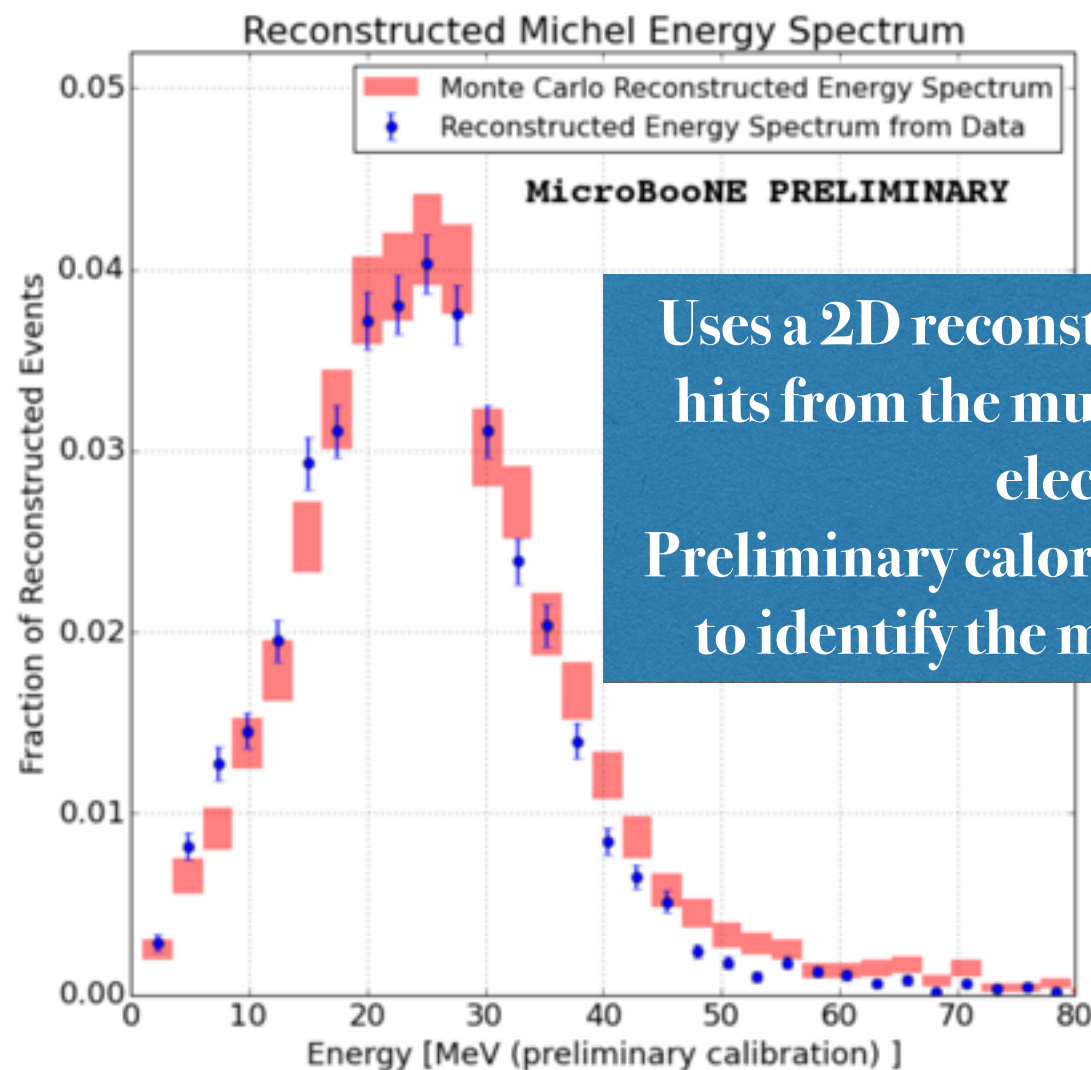
Data/MC comparison for  
Cosmic events

MICROBOONE-NOTE-1014-PUB



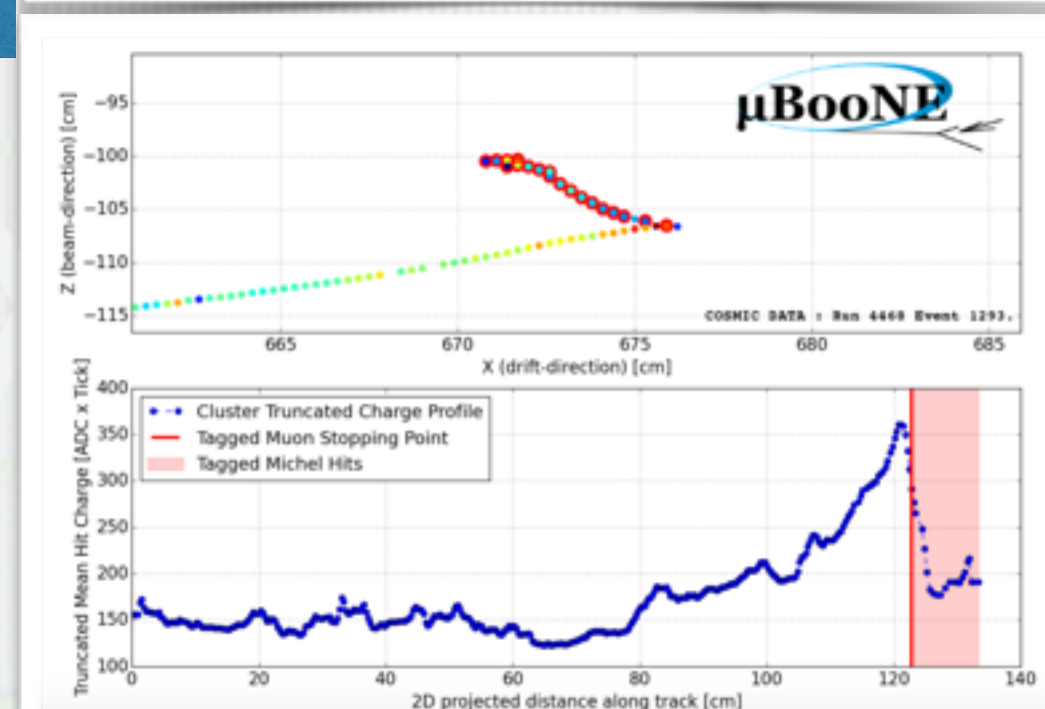
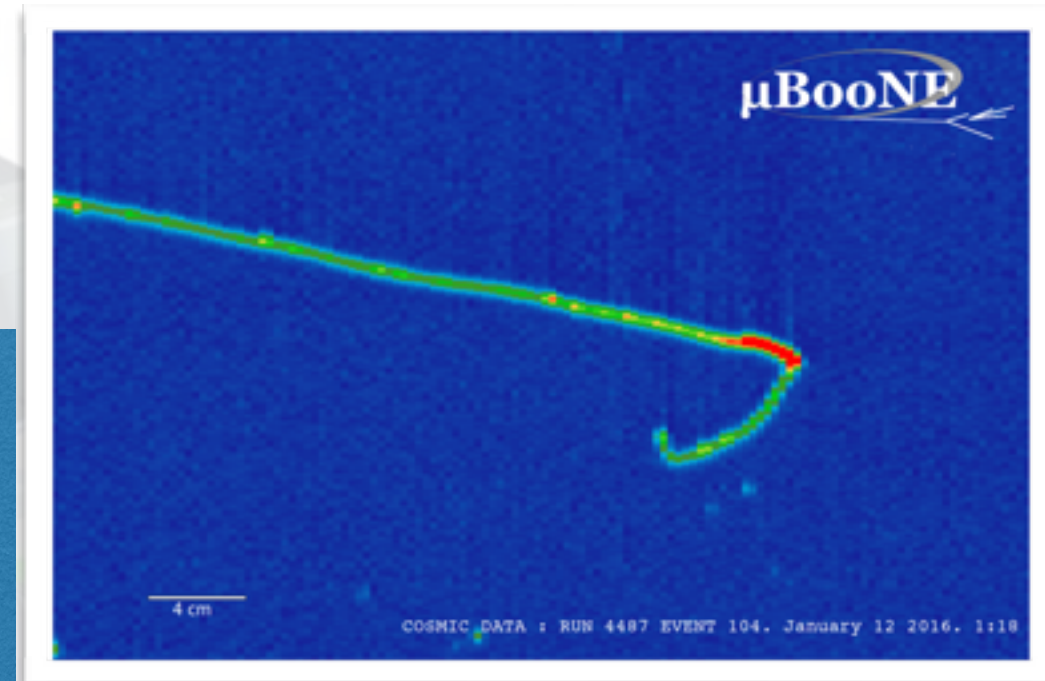
# MicroBooNE Michel Electrons

Studies done using cosmic events



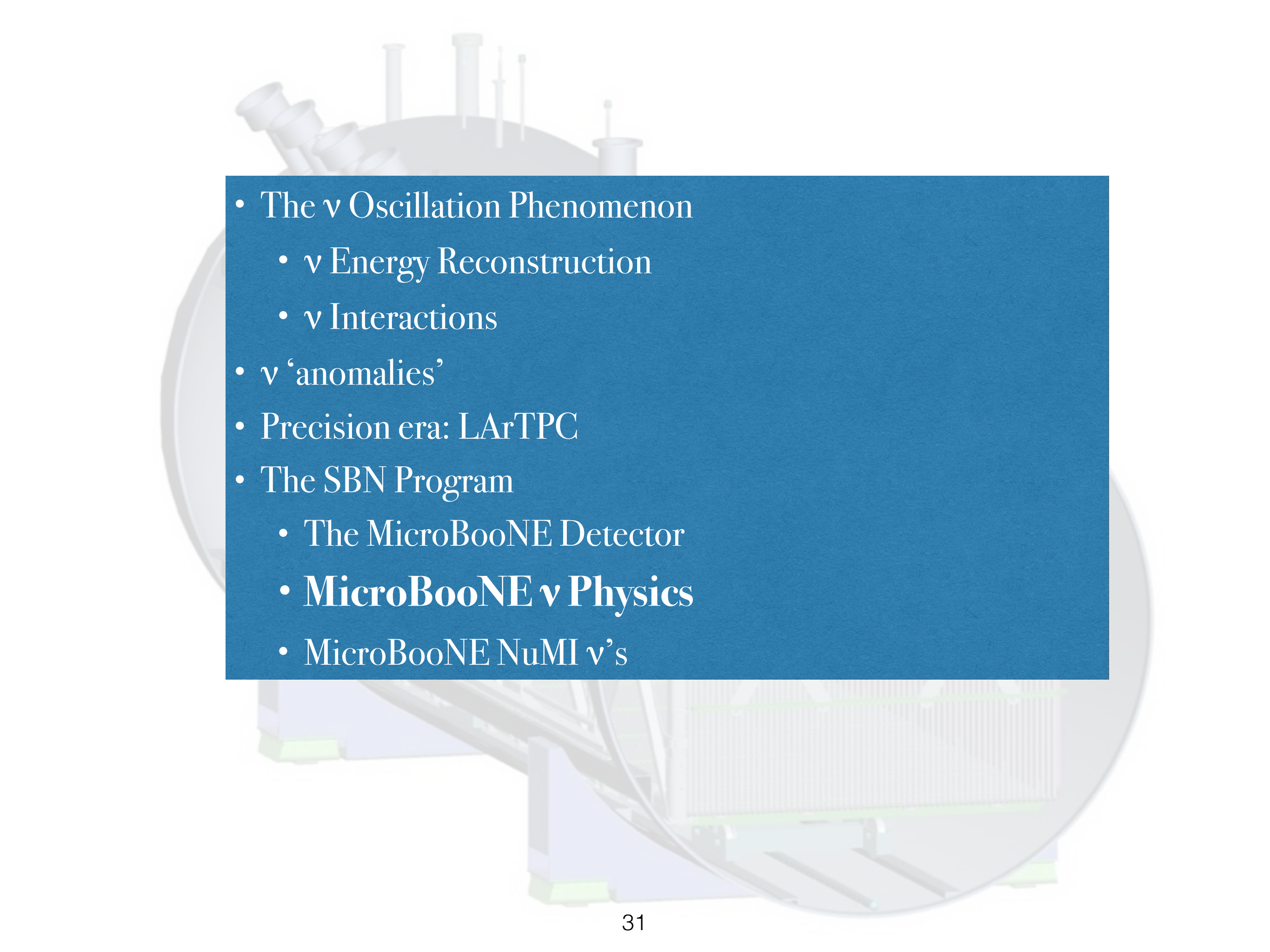
Uses a 2D reconstruction to identify hits from the muon and the Michel electron.

Preliminary calorimetric calibration to identify the muon Bragg peak.



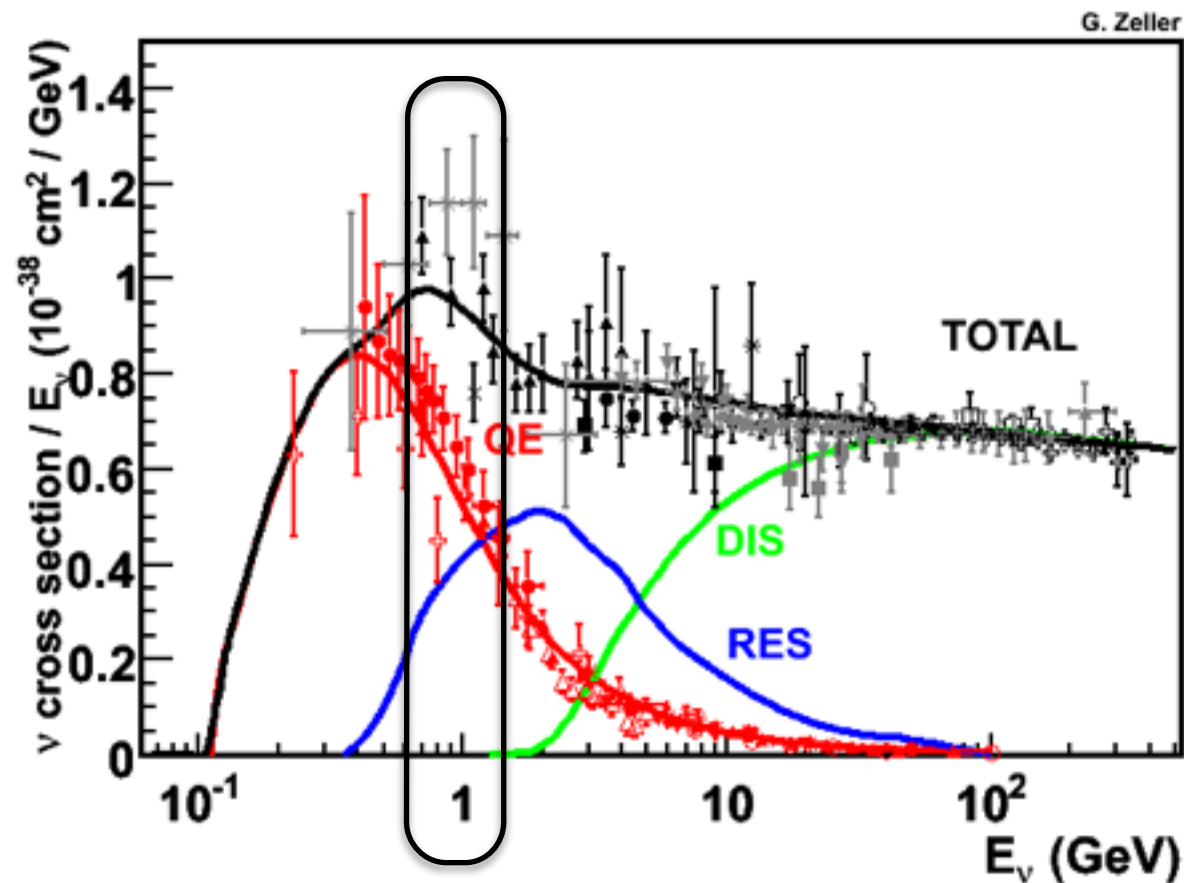
$$E(\text{MeV}) = Q(e^-) \times \frac{23.6}{10^6} \left[ \frac{\text{MeV}}{e^-} \right] \times 1.16[\text{Lifetime Corr.}] \times \frac{1}{0.62}[\text{Recombination Corr.}]$$



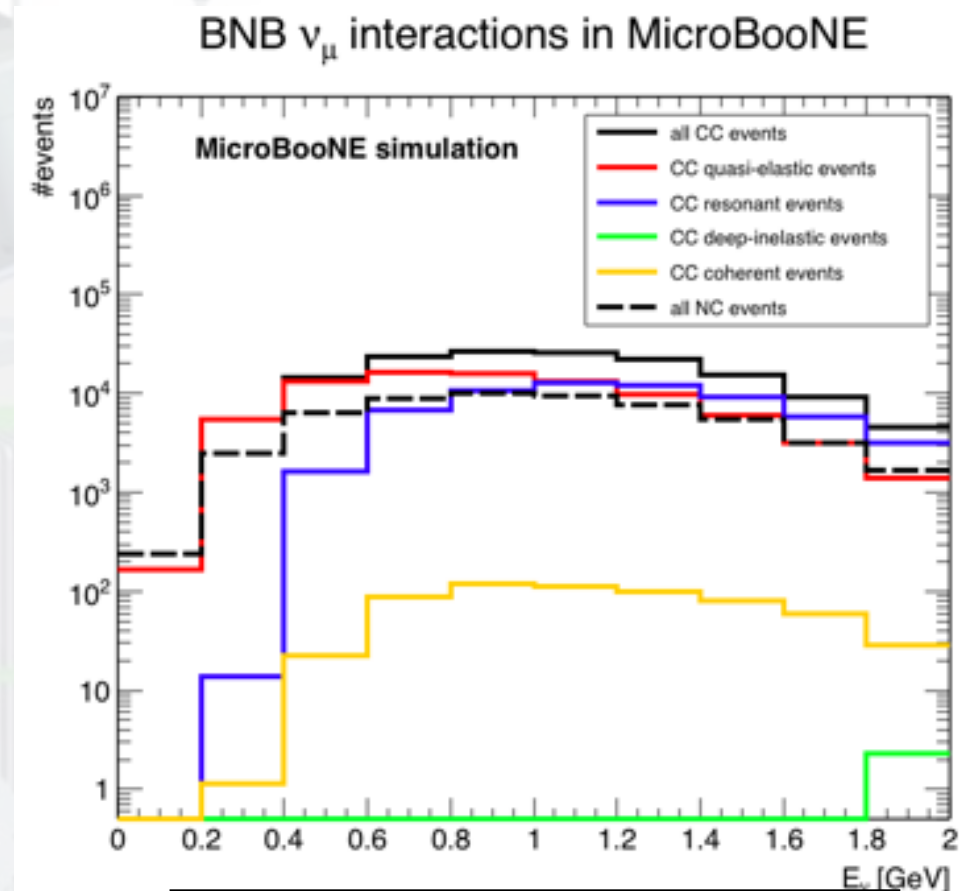
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# MicroBooNE $\nu$ Physics

We need to understand all the possible differences from our data to the current simulations



Present and future  $\nu$  oscillation experiments cover a region full of reaction thresholds and sparse data.

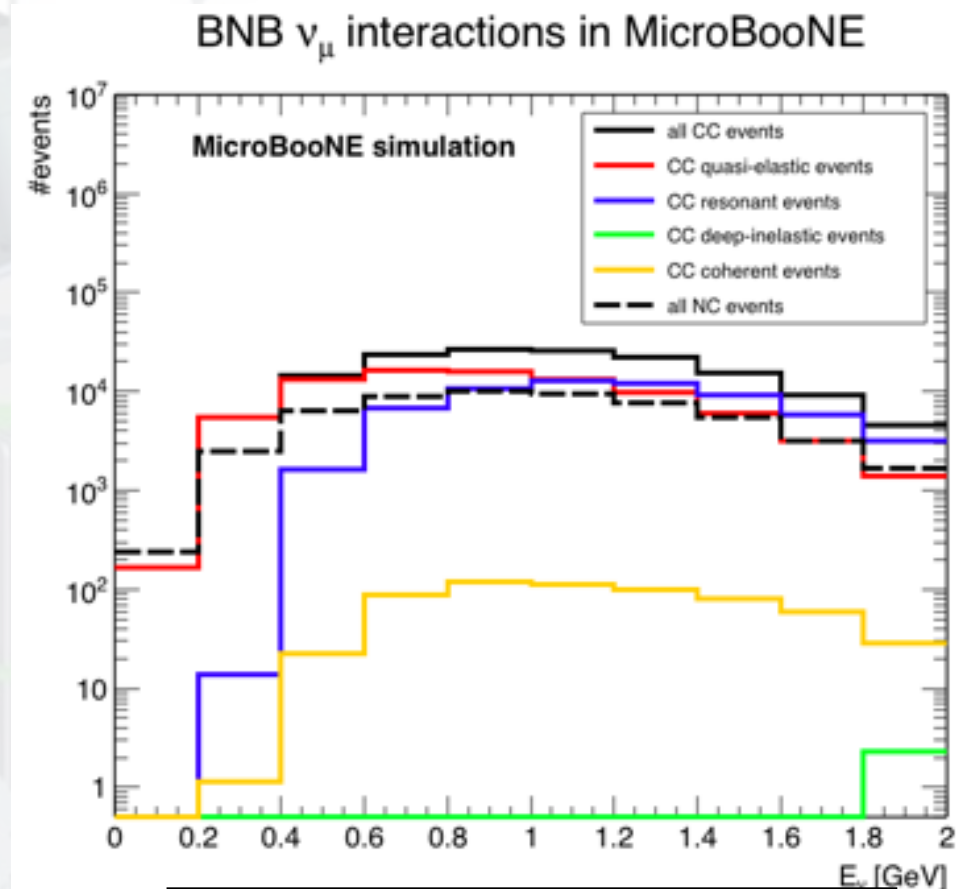
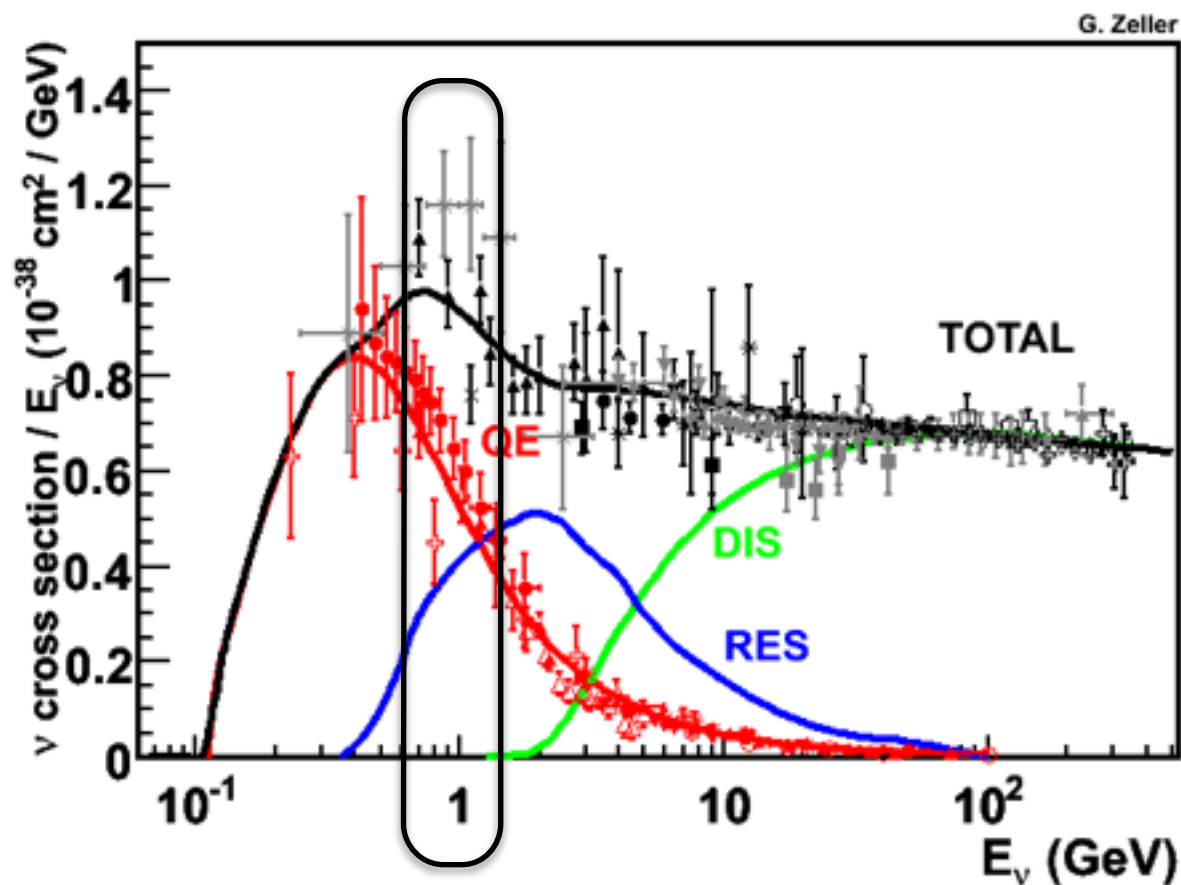


Using GENIE without MEC



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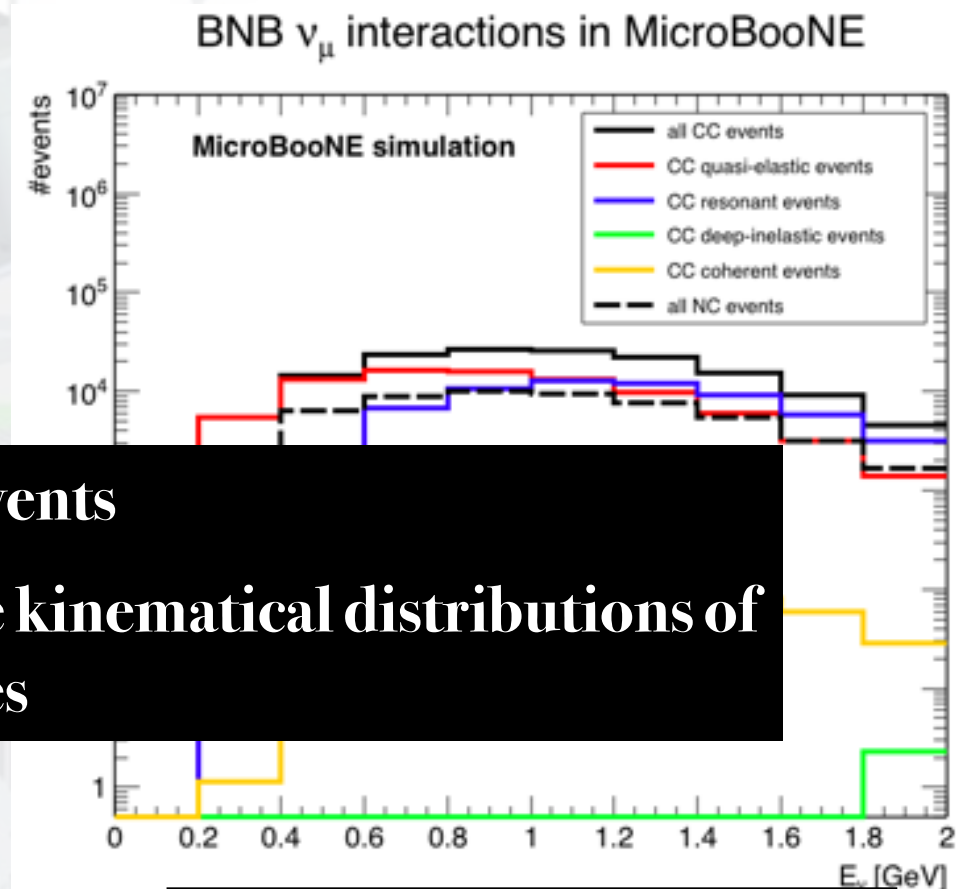
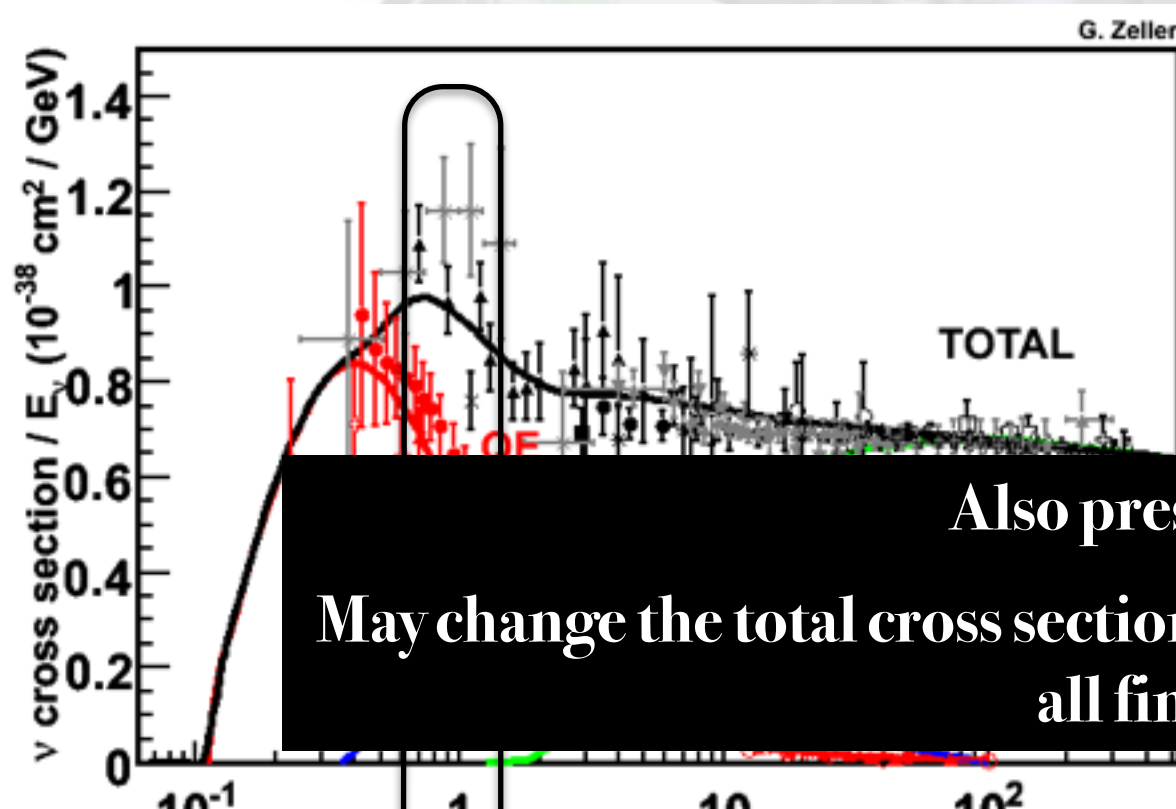
Using GENIE without MEC

Current predictions (GENIE & NuWRO) advance a huge contribution of MEC interactions (20% in  $\nu_{\mu}$  CC).

Even if resonant  $\pi$  production and MEC interactions are big contributions, due to FSI we expect to see mostly  $\text{CC}0\pi^{+/-/0}$  events.

# MicroBooNE $\nu$ Physics

We need to understand all the possible differences from our data to the current simulations



Also present on the  $\nu_e$  events

May change the total cross section & shape of the kinematical distributions of all final state particles

Particularly important at low  $Q^2$  interactions

Using GENIE without MEC

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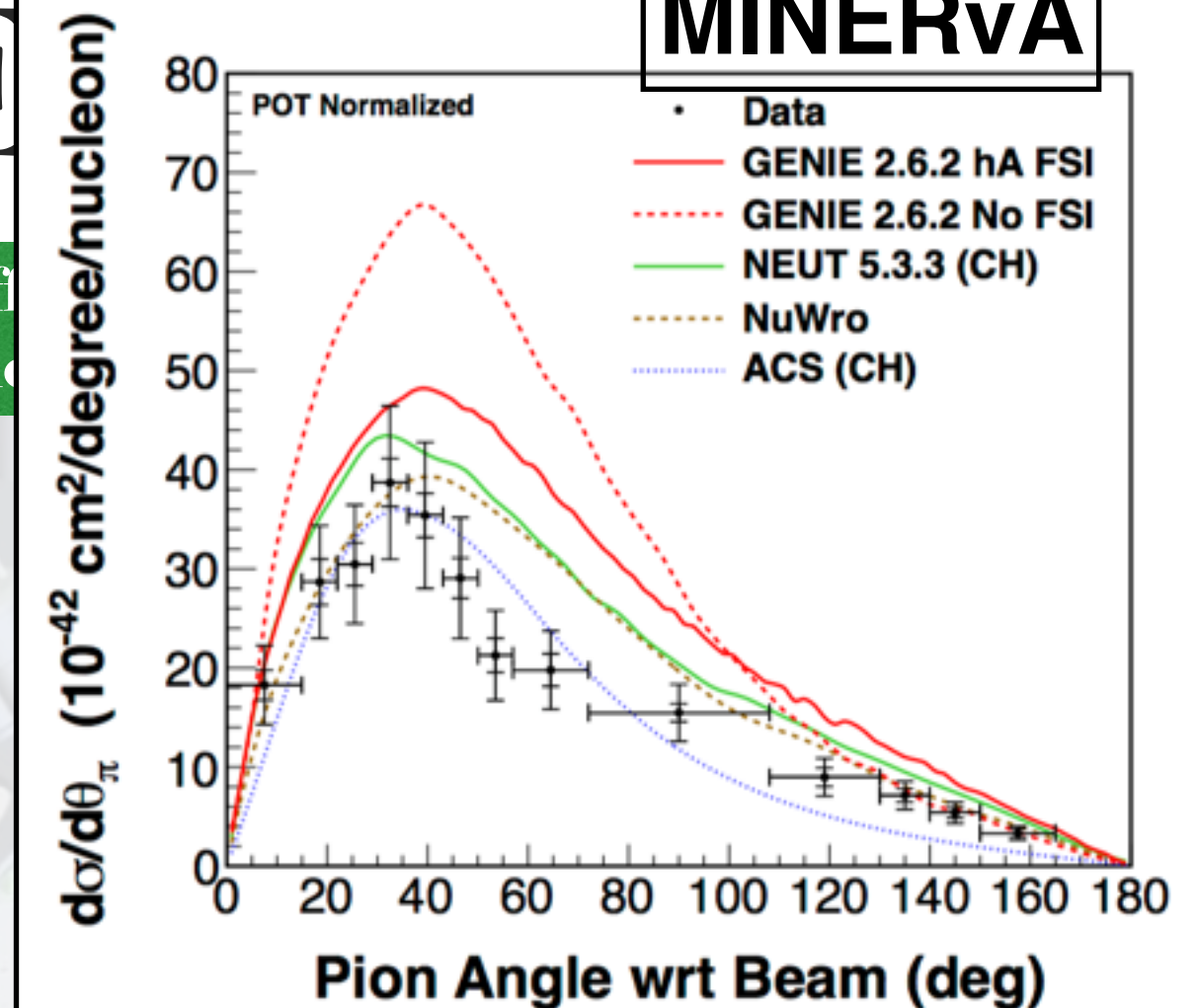
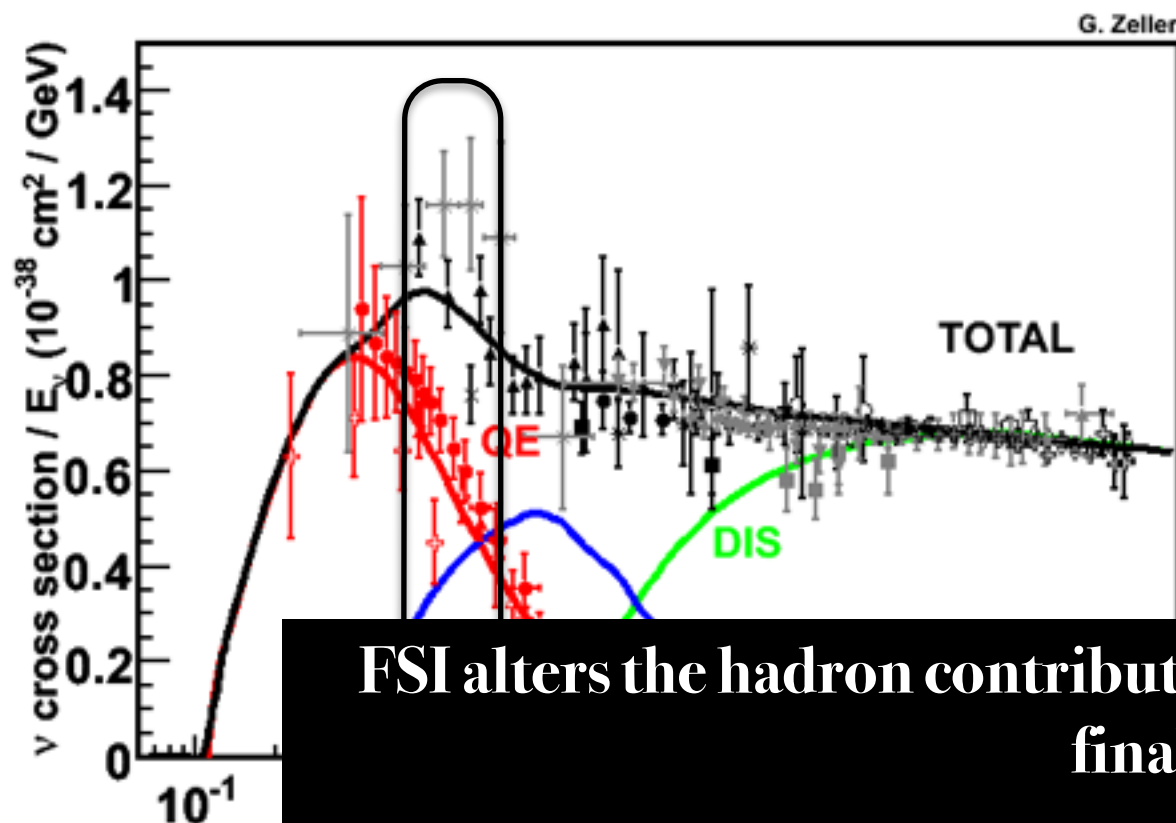
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# MicroBooNE

We need to understand all the possible different simulation

MINERvA



FSI alters the hadron contribution and the shape of the kinematics for the final state hadrons:

reconstructed visible energy is altered

Using GENIE without MEC

Increases difficulty on separate resonant (RO) and a huge contribution of MEC interactions vs MEC (20% in  $\mu \text{ CC}$ ).

Even if resonant pion production and MEC interactions are big contributions, due to FSI we expect to see mostly  $\text{CC}0\pi^{+/-/0}$  events.

# MicroBooNE $\nu$ Physics

First efforts on the  $\nu_\mu$  CC selection for the BNB events

Inclusive  $\nu_\mu$  CC measurements motivation:

- Small theoretical bias (for the lepton kinematics).
- More robust channel for theorists to understand our data.
- No additional difficulties to compare along experiments.
  - This is a topology on what we all of us agree.
- Safer way, using double differential  $(p_\mu, \theta_\mu)$ , to distinguish different neutrino reaction regions (CCQE and resonant).



# MicroBooNE $\nu$ Physics

First efforts on the  $\nu_\mu$  CC selection for the BNB events

## Selection I

- Longest track fully contained ( $\mu$  candidate)
- Candidate  $\mu$  length  $> 75$  cm

## Selection II

- $N=1$ :
  - fully contained ( $\mu$  candidate)
  - length  $\mu > 40$  cm
- $N>1$ :
  - longest track is the  $\mu$  candidate
  - min. length any other track  $> 15$  cm

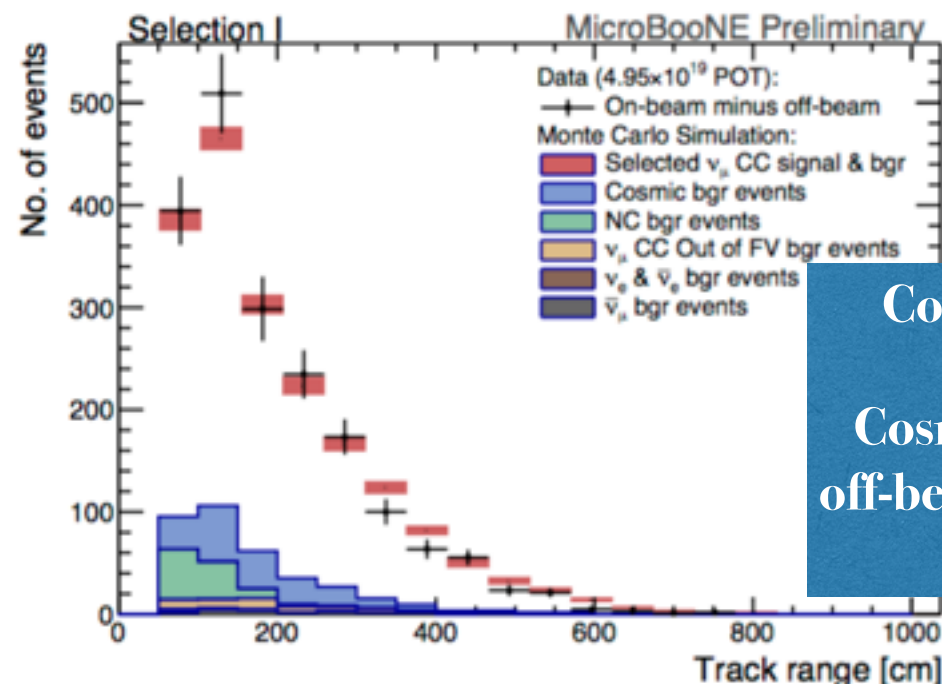
# MicroBooNE $\nu$ Physics

First presented at  
NEUTRINO2016

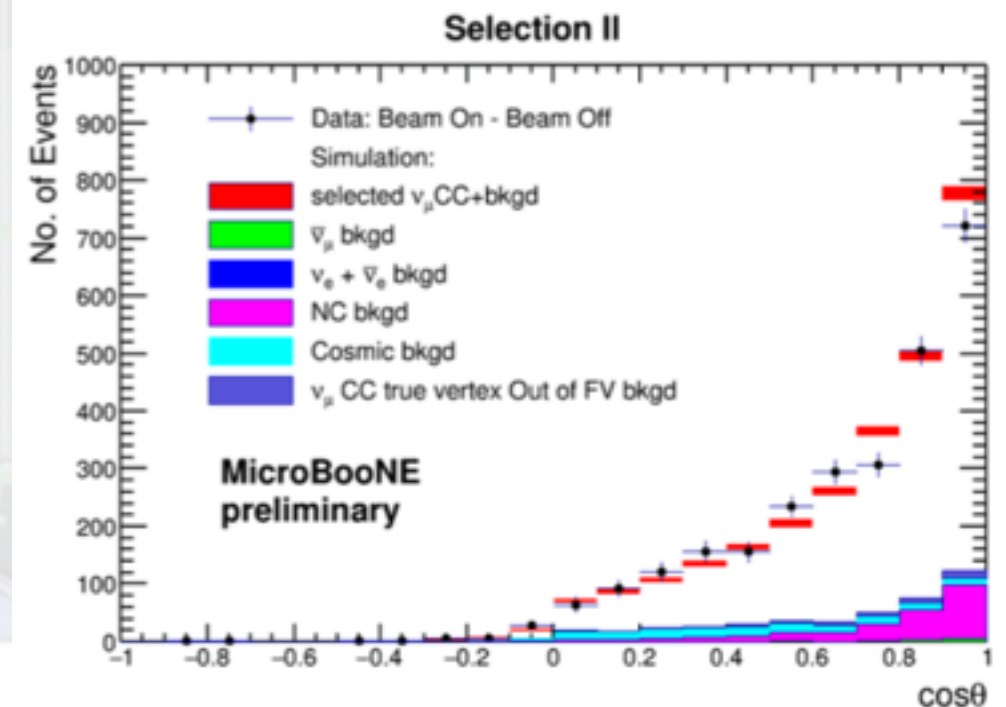
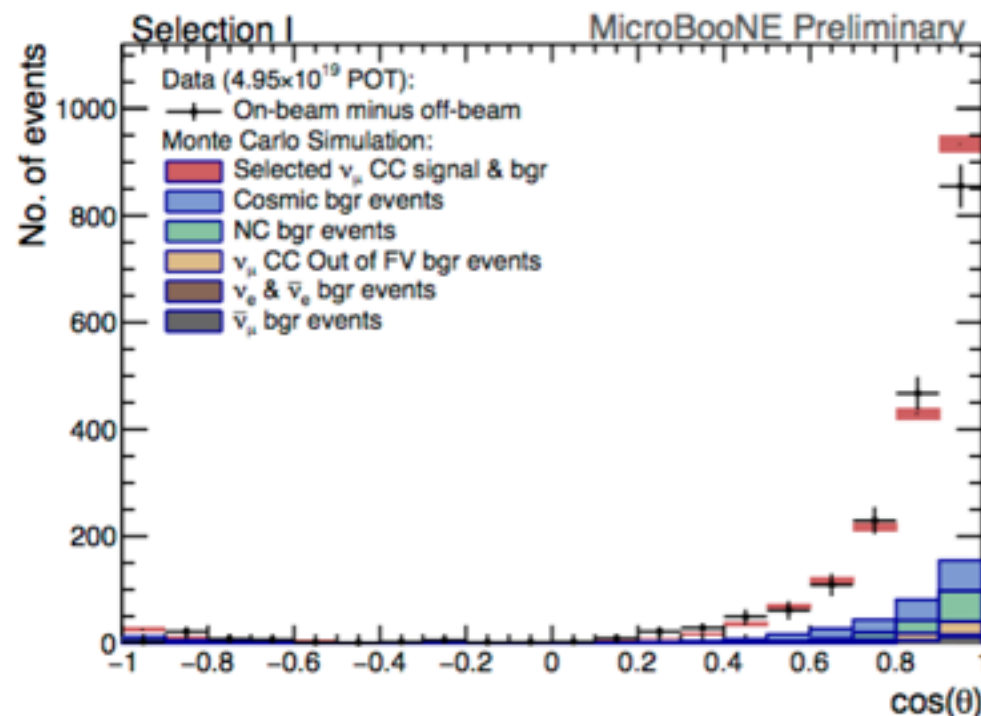
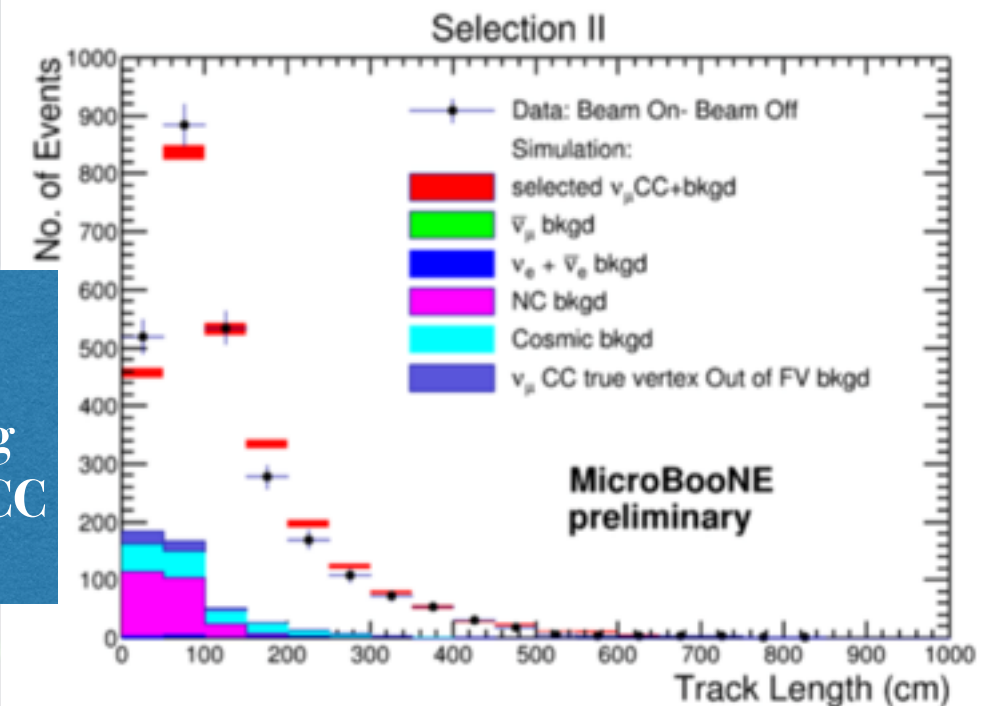
First efforts on the  $\nu_\mu$  CC selection for the BNB events

Selection I

Selection II



Cosmic contamination  
~50%  
Cosmic subtraction using  
off-beam data (passing  $\nu_\mu$  CC  
selection)





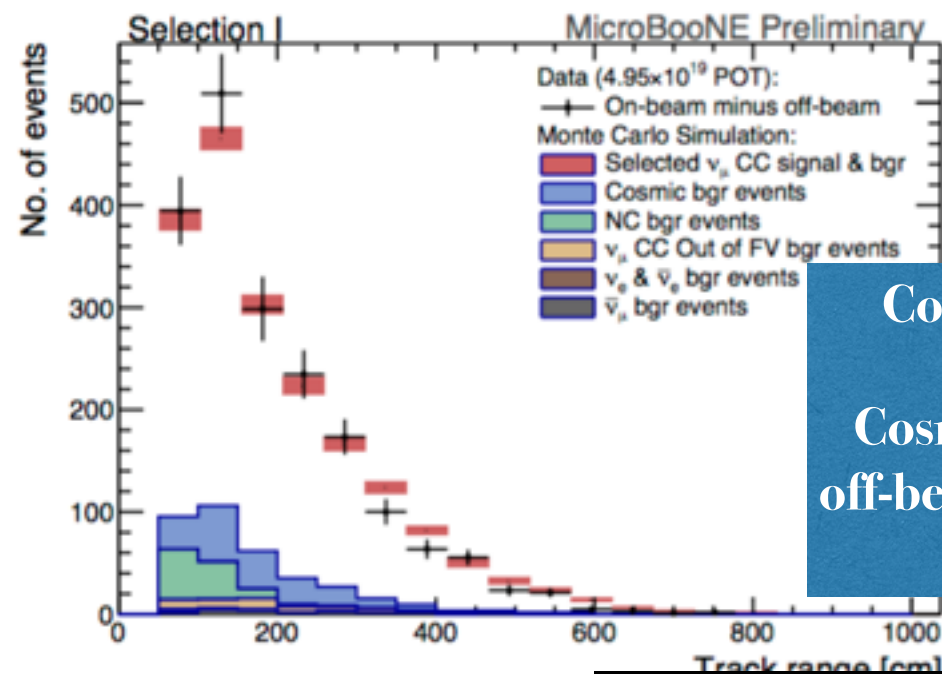
# MicroBooNE $\nu$ Physics

First presented at  
NEUTRINO2016

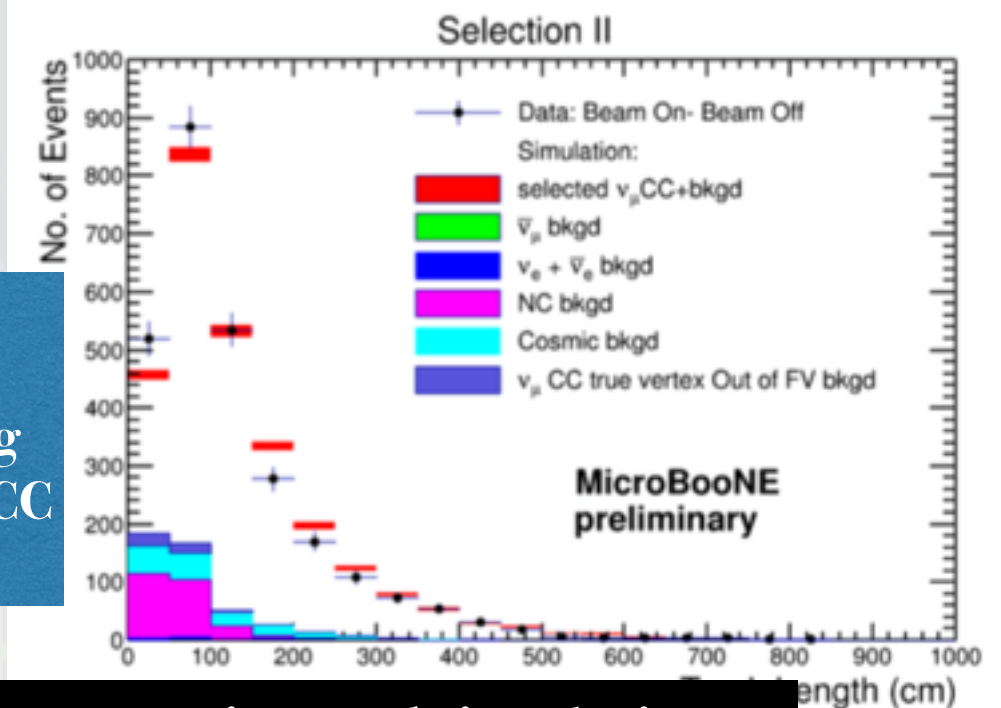
First efforts on the  $\nu_\mu$  CC selection for the BNB events

Selection I

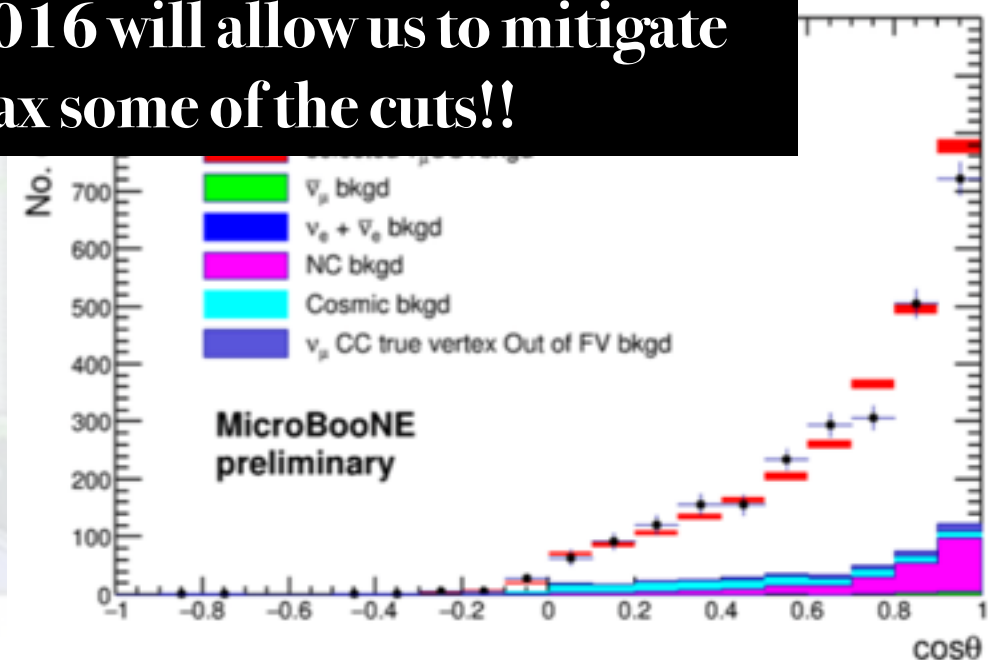
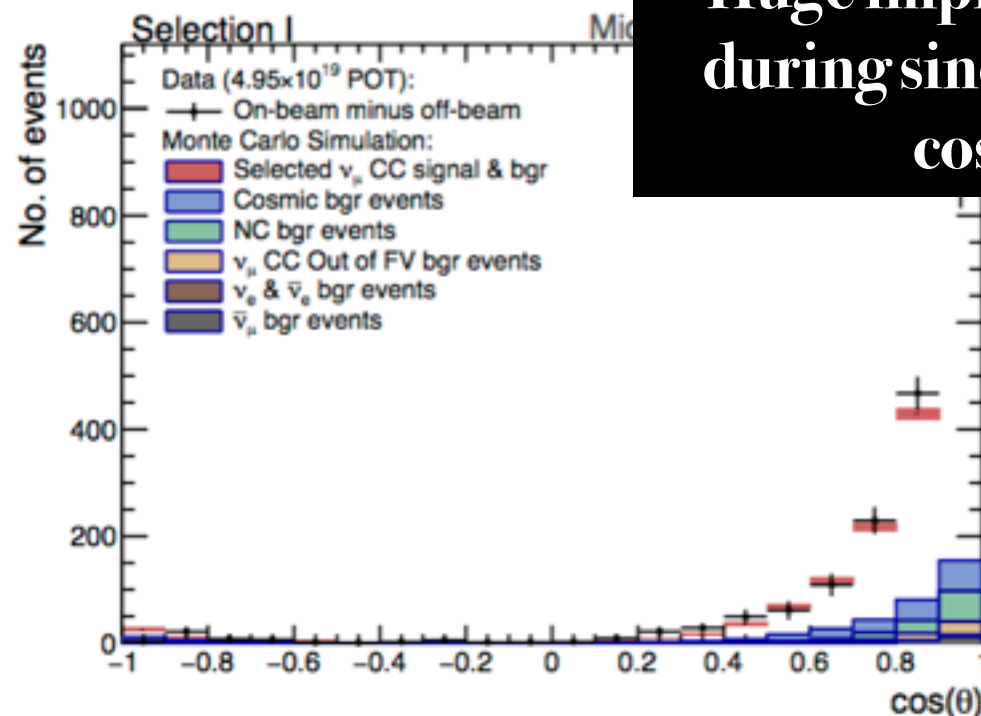
Selection II



Cosmic contamination  
~50%  
Cosmic subtraction using  
off-beam data (passing  $\nu_\mu$  CC  
selection)



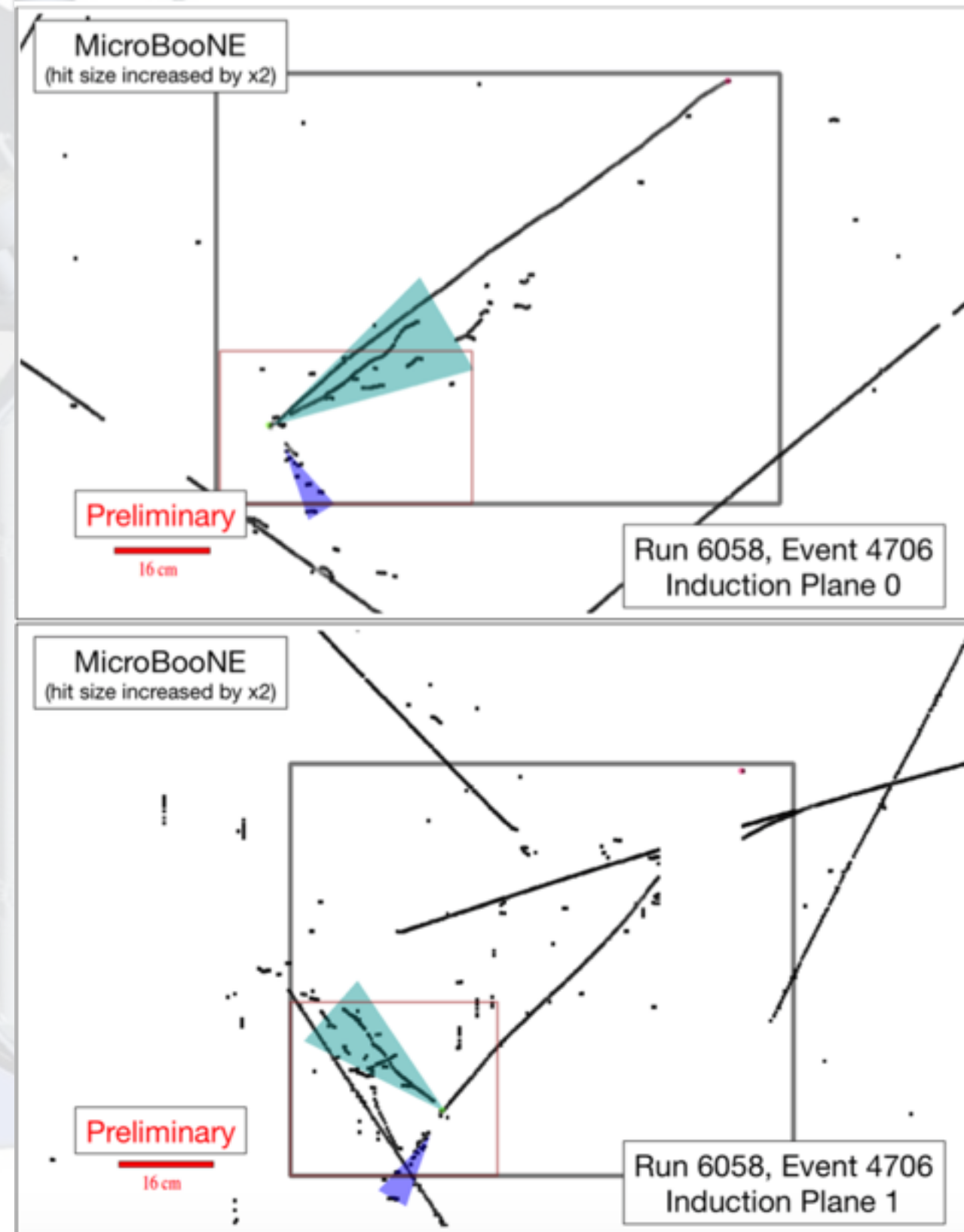
Huge improvements in reconstruction and simulation  
during since NEUTRINO2016 will allow us to mitigate  
cosmic bkg and relax some of the cuts!!



# MicroBooNE $\nu$ Physics

First efforts on the  $\nu_\mu$   $CC\pi^0$  studies

- $CC\pi^0$  are one of the main bkg for  $\nu_e$ .
- These cross section measurements showed poor prediction by the different MC wrt all available experimental data.
- Allow us to train our reconstruction algorithms to improve our shower efficiency and energy resolution.

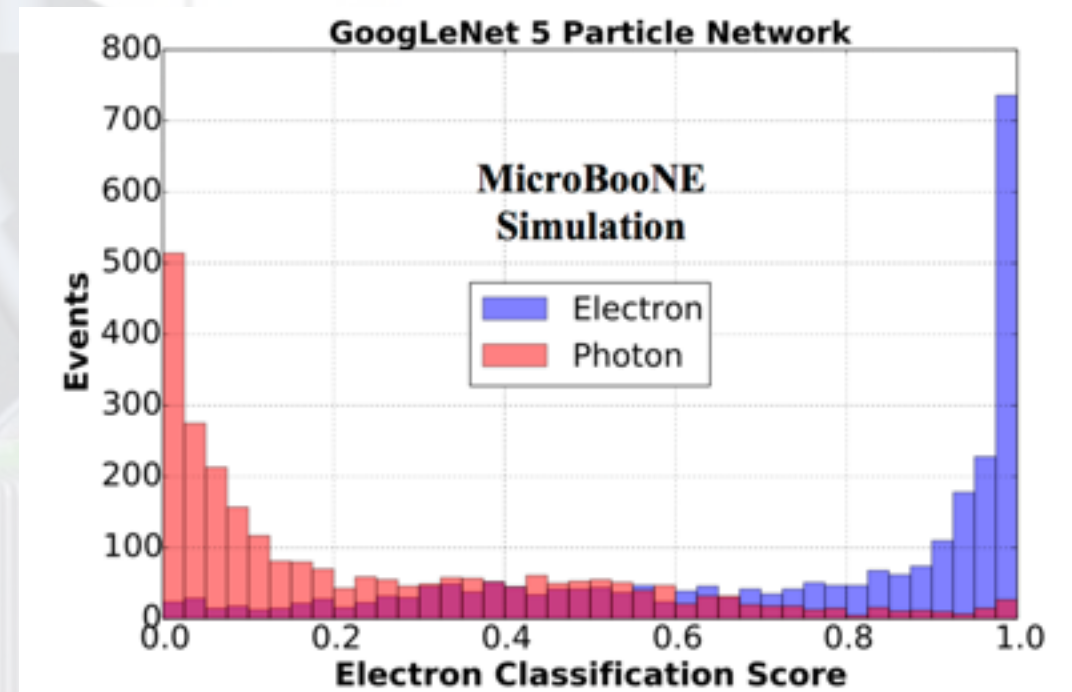
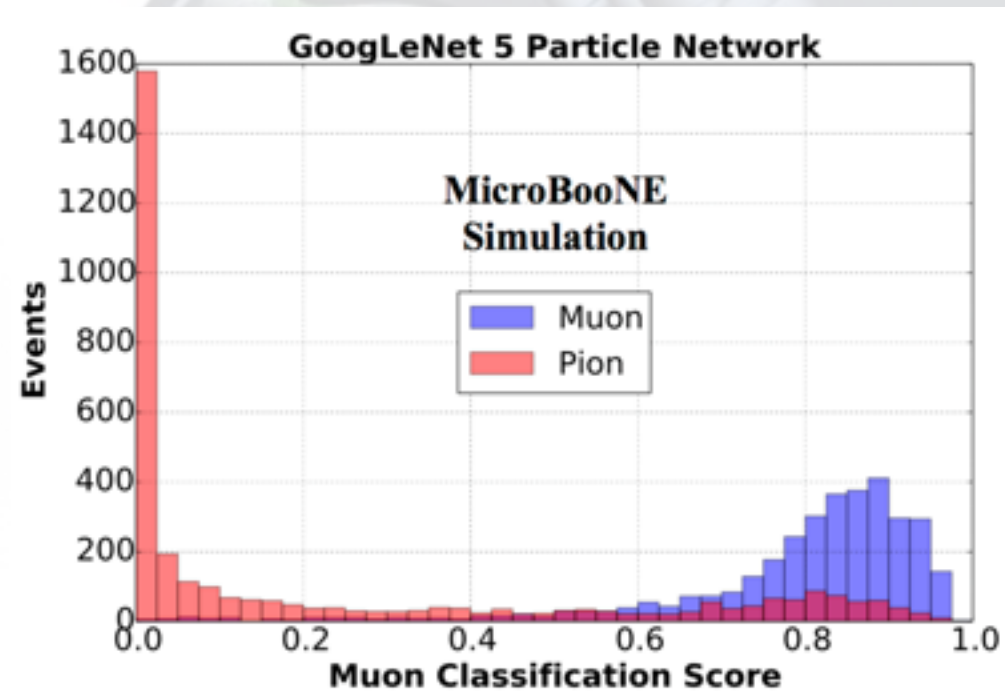


MICROBOONE-NOTE-1012-PUB



# MicroBooNE $\nu$ Physics

Exploring different options to select our neutrinos



First MicroBooNE publication:  
**Convolutional Neural Networks Applied to Neutrino Events in a Liquid Argon  
Time Projection Chamber, arXiv:1611.05531**

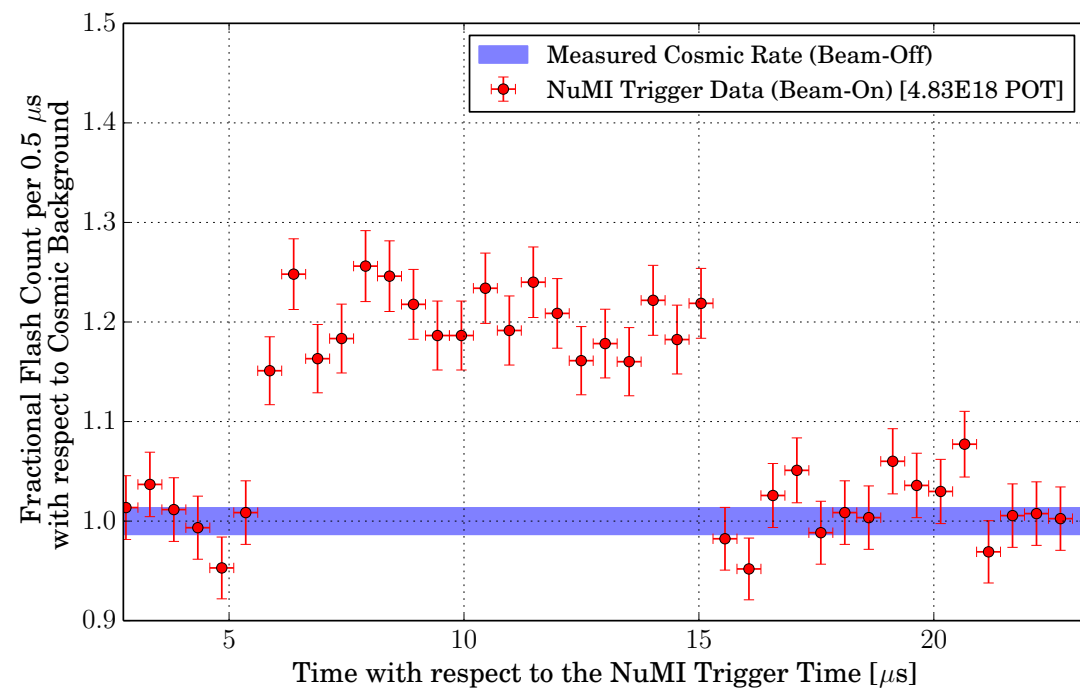
Using Deep Learning techniques to identify  $\nu$  events from cosmics

- 
- The  $\nu$  Oscillation Phenomenon
    - $\nu$  Energy Reconstruction
    - $\nu$  Interactions
  - $\nu$  ‘anomalies’
  - Precision era: LArTPC
  - The SBN Program
    - The MicroBooNE Detector
    - MicroBooNE  $\nu$  Physics
      - **MicroBooNE NuMI  $\nu$ 's**



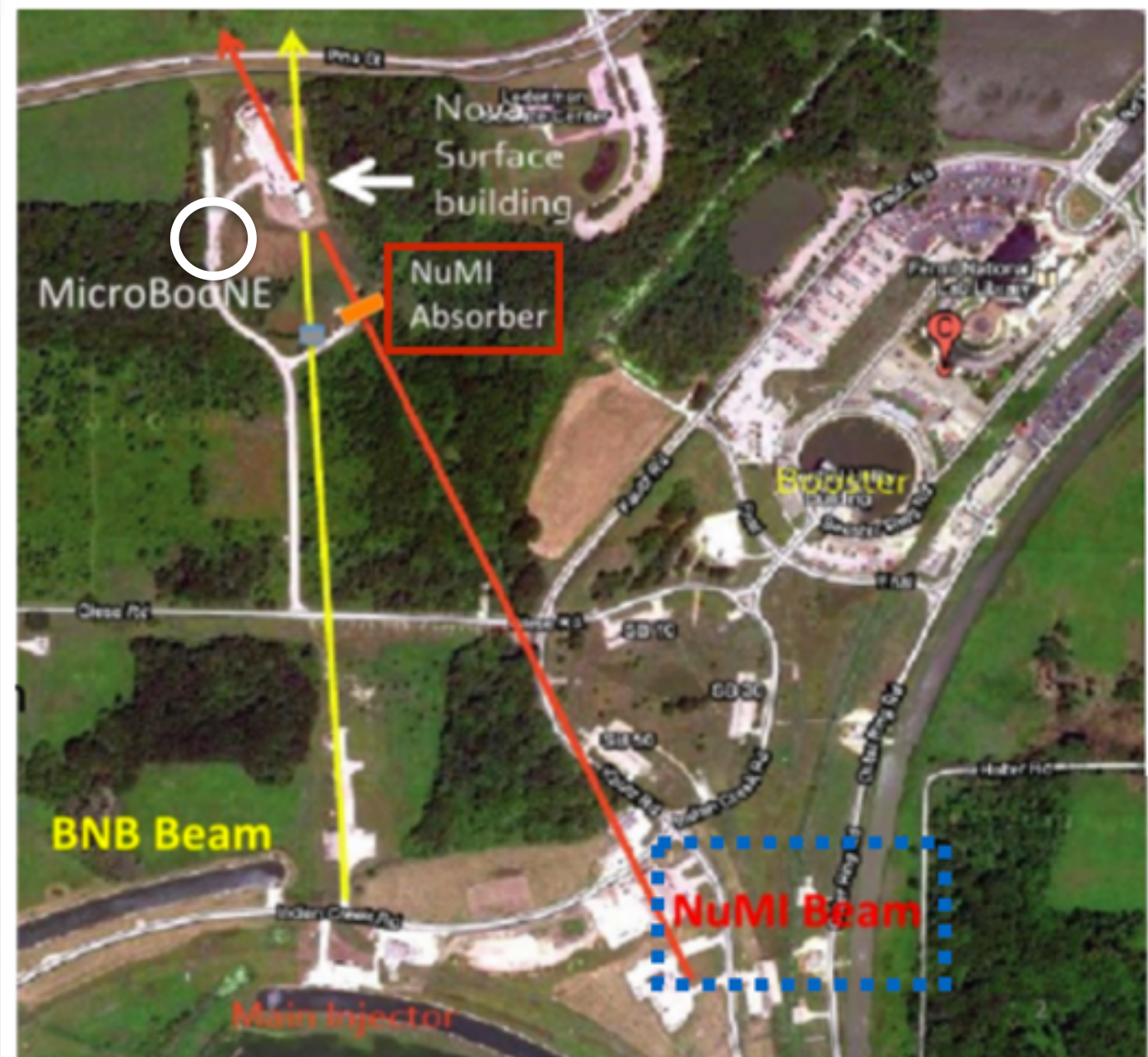
# MicroBooNE $\nu$ Physics

NuMI: On the way to constrain  $\nu$ -Argon interactions



MicroBooNE receives 2 different neutrino beams: BNB but also NuMI  
NuMI arrives off-axis  $\sim 6$  degrees wrt  $z$

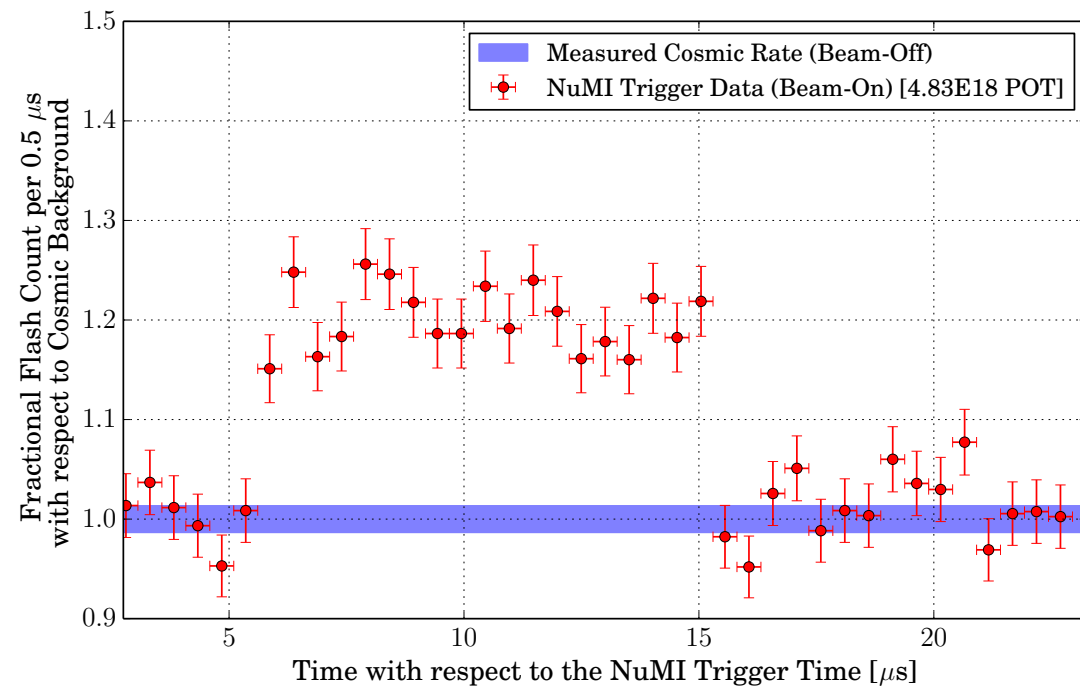
Double the beam, double the fun!





# MicroBooNE $\nu$ Physics

## NuMI: On the way to constrain $\nu$ -Argon interactions



Ongoing studies on  $\nu_e$  CC inclusive cross section.

- Fundamental to understand the  $\nu_e$  spectrum in a non-oscillation hypothesis

Ongoing different studies in  $\nu_\mu$  events.

- Fundamental to have a first comprehension of  $\nu$ -Ar interactions and their uncertainties
- Understand  $\nu_e / \nu_\mu$

Fundamental to understand if any possible anomaly in the BNB induced neutrino interactions are due to:

- lepton differences
- cross section unknowns

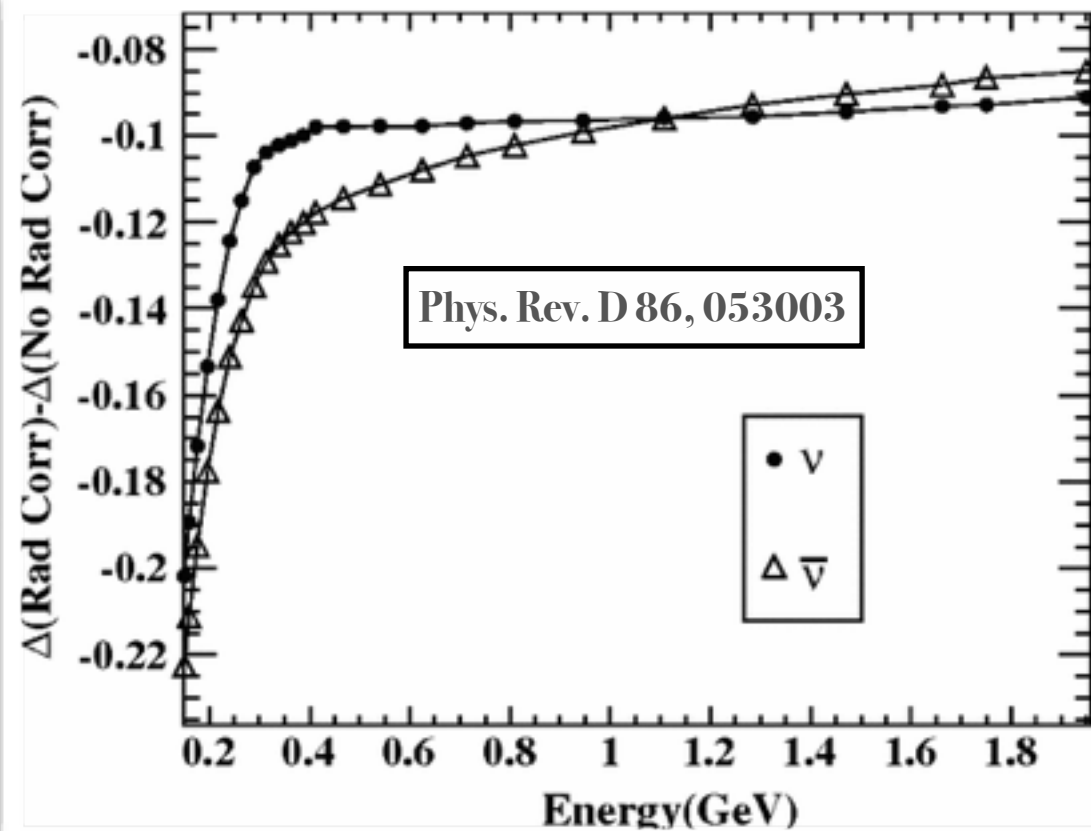
NuMI has been producing neutrino data for MicroBooNE during 1.5 year and now it is producing anti-neutrino data.

- This will allow us to start studies  $\nu$ /anti- $\nu$



# MicroBooNE $\nu$ Physics

NuMI: On the way to constrain  $\nu$ -Argon interactions



Calculations shows differences in  $\nu_e / \nu_\mu$  induced by:

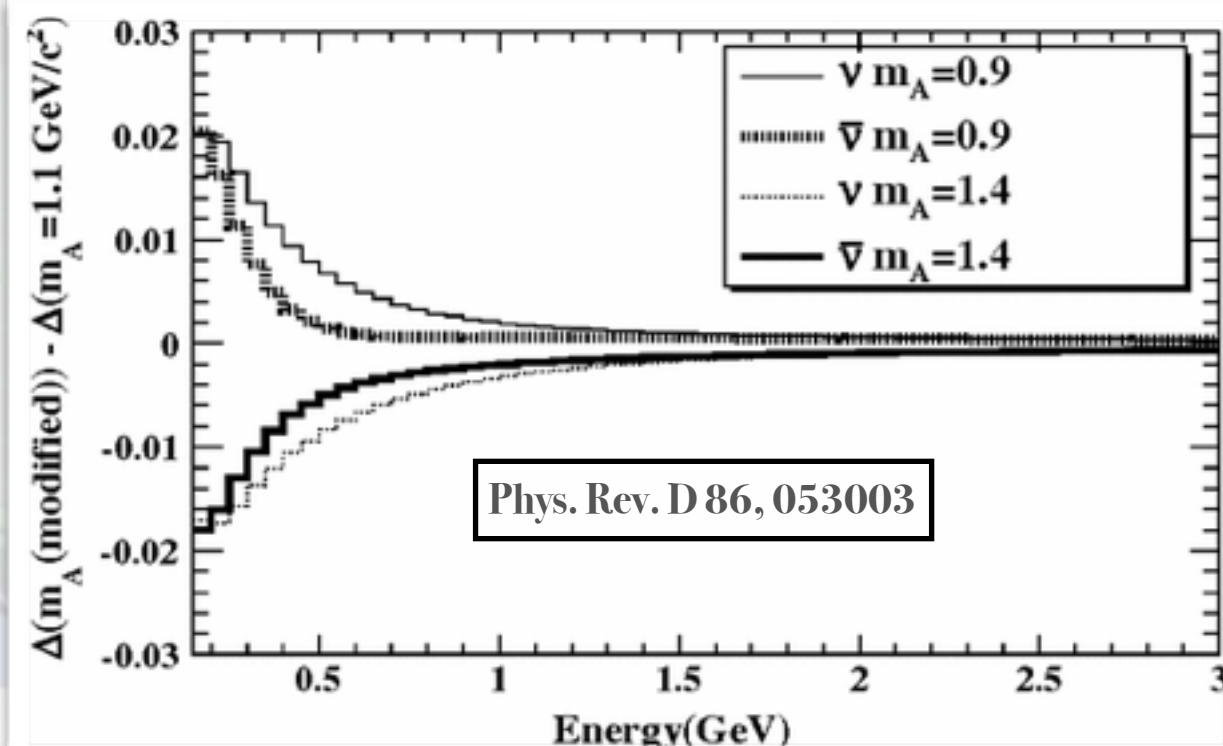
- form factors
- radiative corrections
- lepton mass

Ongoing studies on  $\nu_e$  CC inclusive cross section.

- Fundamental to understand the  $\nu_e$  spectrum in a non-oscillation hypothesis

Ongoing different studies in  $\nu_\mu$  events.

- Fundamental to have a first comprehension of  $\nu$ -Ar interactions and their uncertainties
- Understand  $\nu_e / \nu_\mu$



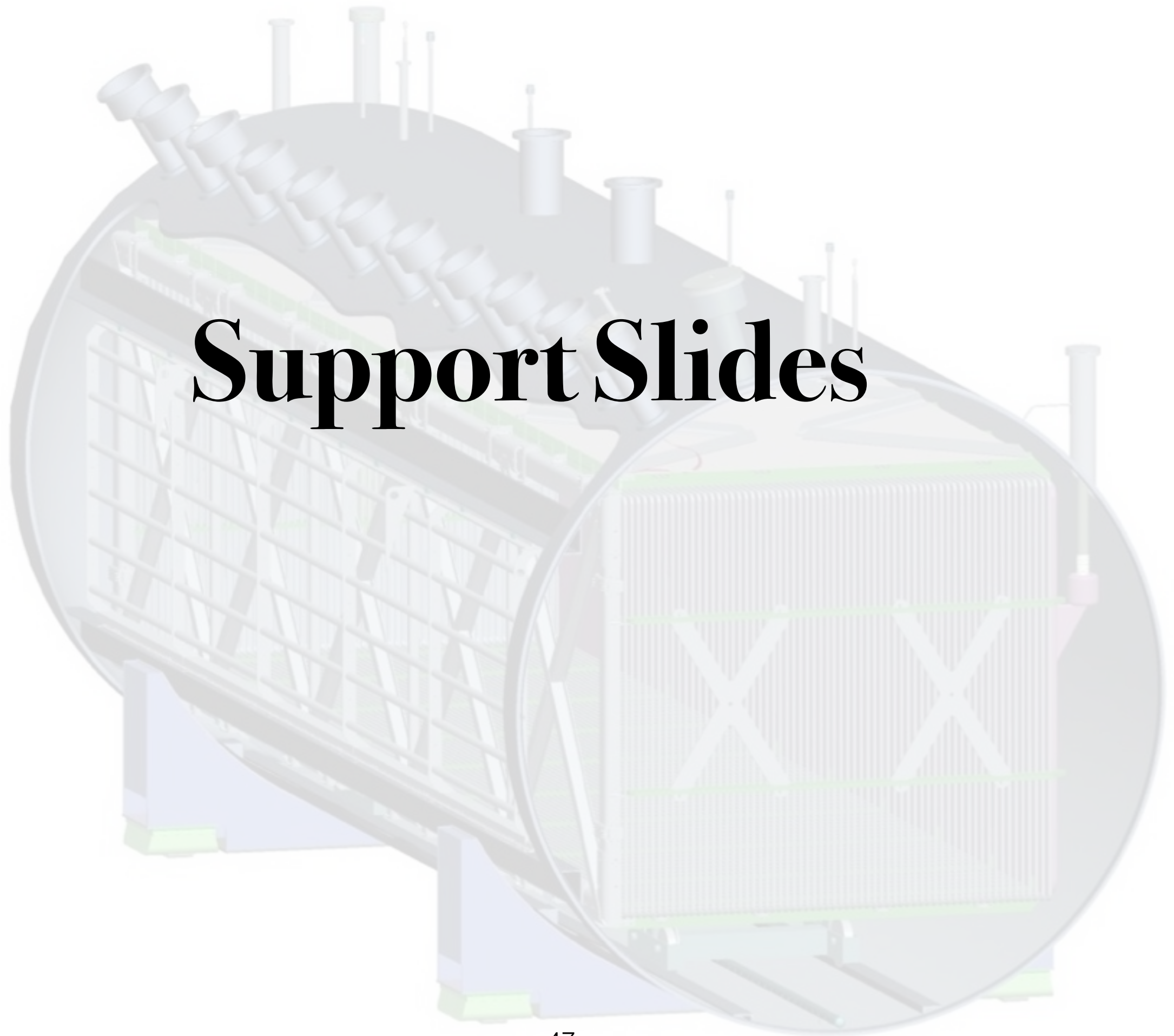
# Conclusions

- Exploring anomalies seen by other experiments. Different possible explanations:
  - $\nu$  cross section mis-modeling/uncertainties ( $\pi$  background unknowns, lepton differences, nuclear effects)
  - Non-Standard Physics (e.g. sterile and heavy neutrinos)
- LArTPC detectors can achieve unprecedented precision in:
  - $e/\gamma$  separation
  - Biggest phase-space covered since deuterium experiments
- MicroBooNE is taking neutrino data since more than 1 year. First large LArTPC operating with front end cold electronics
- Ongoing a huge effort to characterize the detector. Developing new tools for track and shower 3D reconstruction
- Working on produce first  $\nu$ -Ar cross sections at low energies
- Taking data with 2 different  $\nu$  beams
- Part of the SBN program

Lot of physics to come!!!



# Support Slides



# The $\nu$ Oscillation Phenomenon

If we don't have a monochromatic neutrino beam we have to reduce assumptions  
in our extrapolations:

Cover maximum phase-space as possible

Produce measurements as less model dependent as possible to avoid assumptions  
on the different energy ranges

Accurate precision

Accuracy on neutrino  
performance allow  
oscillations with h

We obtain information from  
neutrino events which were  
invisible to us since the  
deuterium era!!!

**Carbon/Water**



**Argon**

Final state particles can be detected  
from 200 MeV

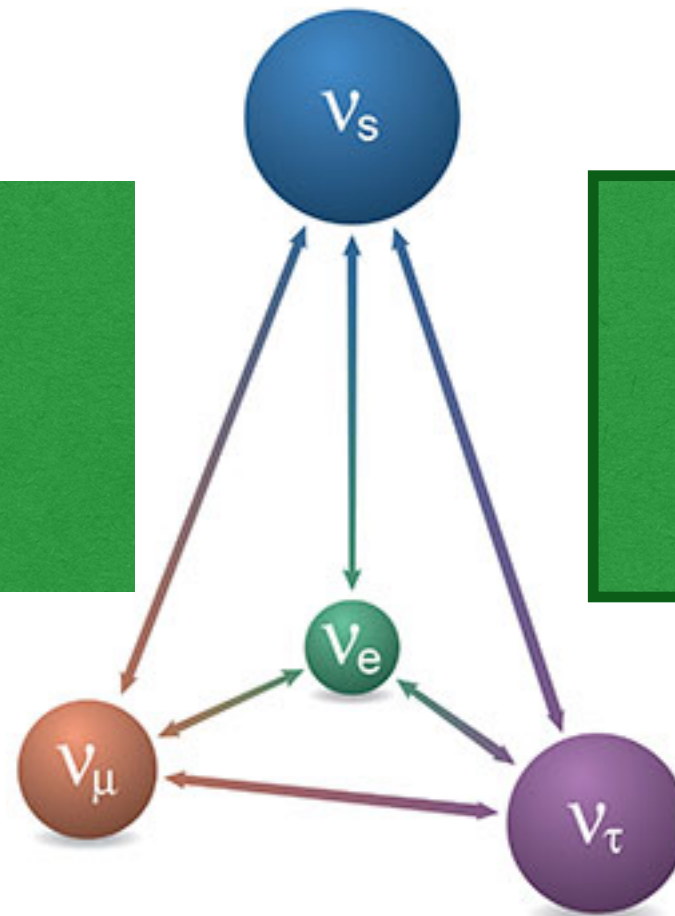
Final state particles can be detected  
from 20 MeV



# Why do we look for sterile $\nu$ ?

## Regular Standard Model $\nu$ :

massive (spin -1/2) particles  
no electric or color charge  
Affected by electroweak force



## Sterile $\nu$ :

massive (spin -1/2) particles  
no electric or color charge  
Not affected by any force

- They can, in principle, be their own anti-particle ('sterile Majorana  $\nu$ ').
- Since they are not associated with either the strong nuclear or electroweak symmetry breaking scales, in principle they can have an arbitrary large/small mass!!
- Since they do have mass and at low energies they act like regular Standard Model Neutrinos, they can participate in neutrino flavor oscillations.



# MicroBooNE $\nu$ Physics

First efforts on the  $\nu_\mu$  CC selection for the BNB events

## Selection I

- Longest track fully contained ( $\mu$  candidate)
- Start position  $\mu$  candidate in FV
- Candidate  $\mu$  start position-tracks within 3cm
- Track and flash must be matched in z
- Candidate  $\mu$  length  $> 75$  cm

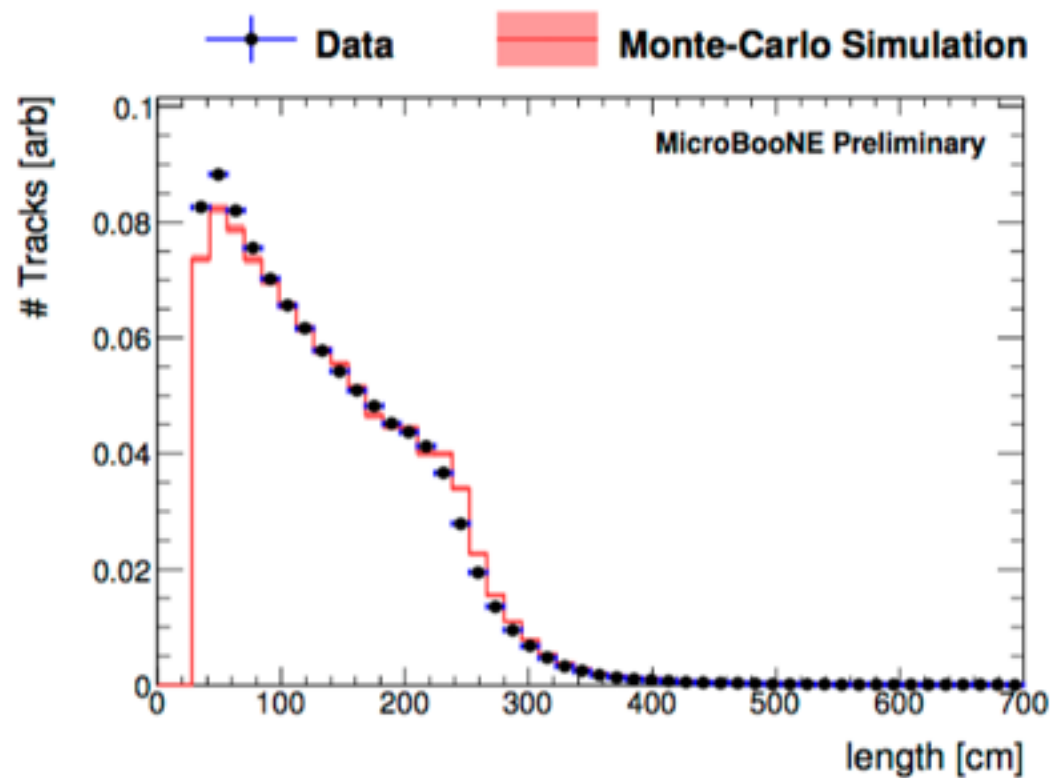
## Selection II

- If only 1 track: fully contained ( $\mu$  candidate)
  - Otherwise, longest track is the  $\mu$  candidate
- Start position  $\mu$  candidate in FV
- Candidate  $\mu$  start position-tracks within 3cm
- Track and flash must be matched in z
- If only 1 track:
  - candidate  $\mu$  length  $> 40$  cm
- If  $> 1$  track,
  - min. length any other track  $> 15$  cm
- Michel removal



# MicroBooNE $\nu$ Physics

We combine hit information from the 3 planes to reconstruct 3D objects



Area normalized

Data/MC comparison for  
neutrino& cosmic events

